

LASER INTERFEROMETER SPACE ANTENNA SCIENCE

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LEIDEN OBSERVATORY
THE NETHERLANDS

MULTIMESSANGERS WINTER WORKSHOP @PRAGUE 4-7 DECEMBER '19

LISA SCIENCE GROUP

Team leader : Jon Gair

*Deputy team leader : **Elena Maria Rossi**, Michele Vallisneri*

An artistic rendering of the LISA mission concept. In the top left corner, a portion of the Earth is visible, showing blue oceans and white clouds. The background is a deep blue space filled with numerous stars and faint nebulae. In the center, a bright, intense white and yellow star is the focal point, with three sharp red laser beams originating from it and extending outwards to three smaller, bright red star-like points, forming a triangular configuration. The text 'LISA' is written in large, white, serif capital letters, and 'Laser Interferometer Space Antenna' is written below it in a smaller, white, serif font.

LISA

Laser Interferometer Space Antenna

A proposal in response to the ESA call for L3 mission concepts

Lead Proposer
Prof. Dr. Karsten Danzmann

- **2013 Gravitational Universe** selected by ESA as L3 science theme within the Cosmic Vision 2015-2025
- **2017 LISA proposal selected** as ESA's L3 mission
- **2034 nominal launch**

The background of the slide is a deep space image featuring a portion of Earth in the upper left corner. The central focus is a bright, multi-colored star (pink, purple, and white) with several sharp red laser beams radiating from it across the dark, star-filled sky. The text 'LISA' is prominently displayed in a large, white, serif font, with 'Laser Interferometer Space Antenna' written below it in a smaller, white, sans-serif font.

LISA

Laser Interferometer Space Antenna

A proposal in response to the ESA call for L3 mission concepts

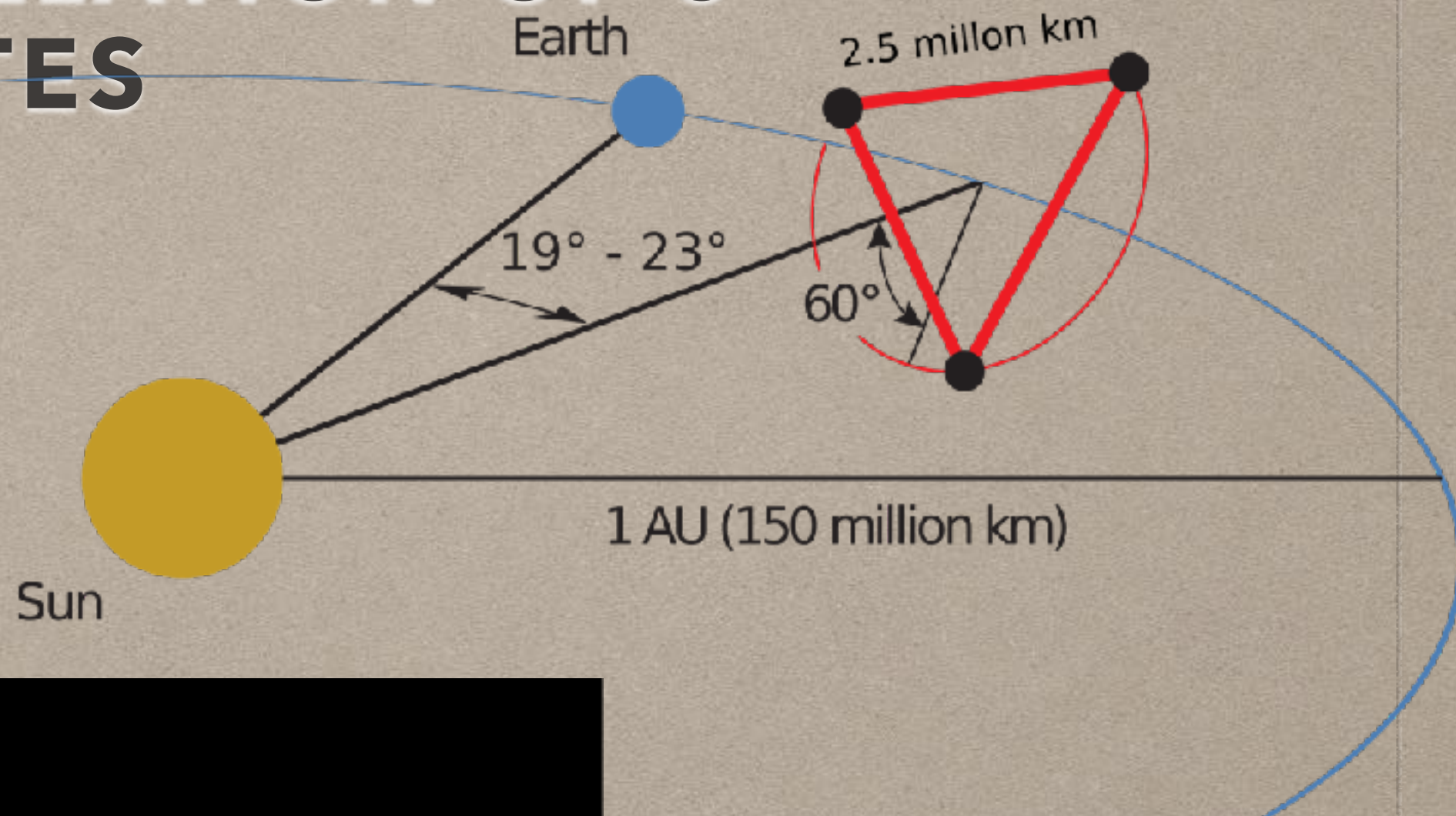
Lead Proposer
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- ~~2034~~ nominal **launch**

News! ESA ministers commit to biggest budget ever for next decay:

Günther Hasinger , ESA
Science Director "**LISA in 2032**"

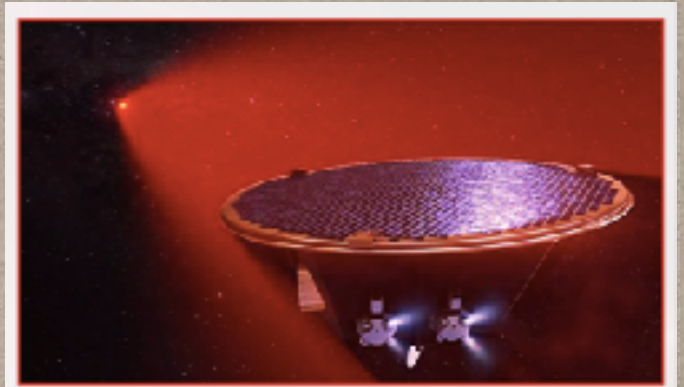
CONSTELLATION OF 3 SATELLITES



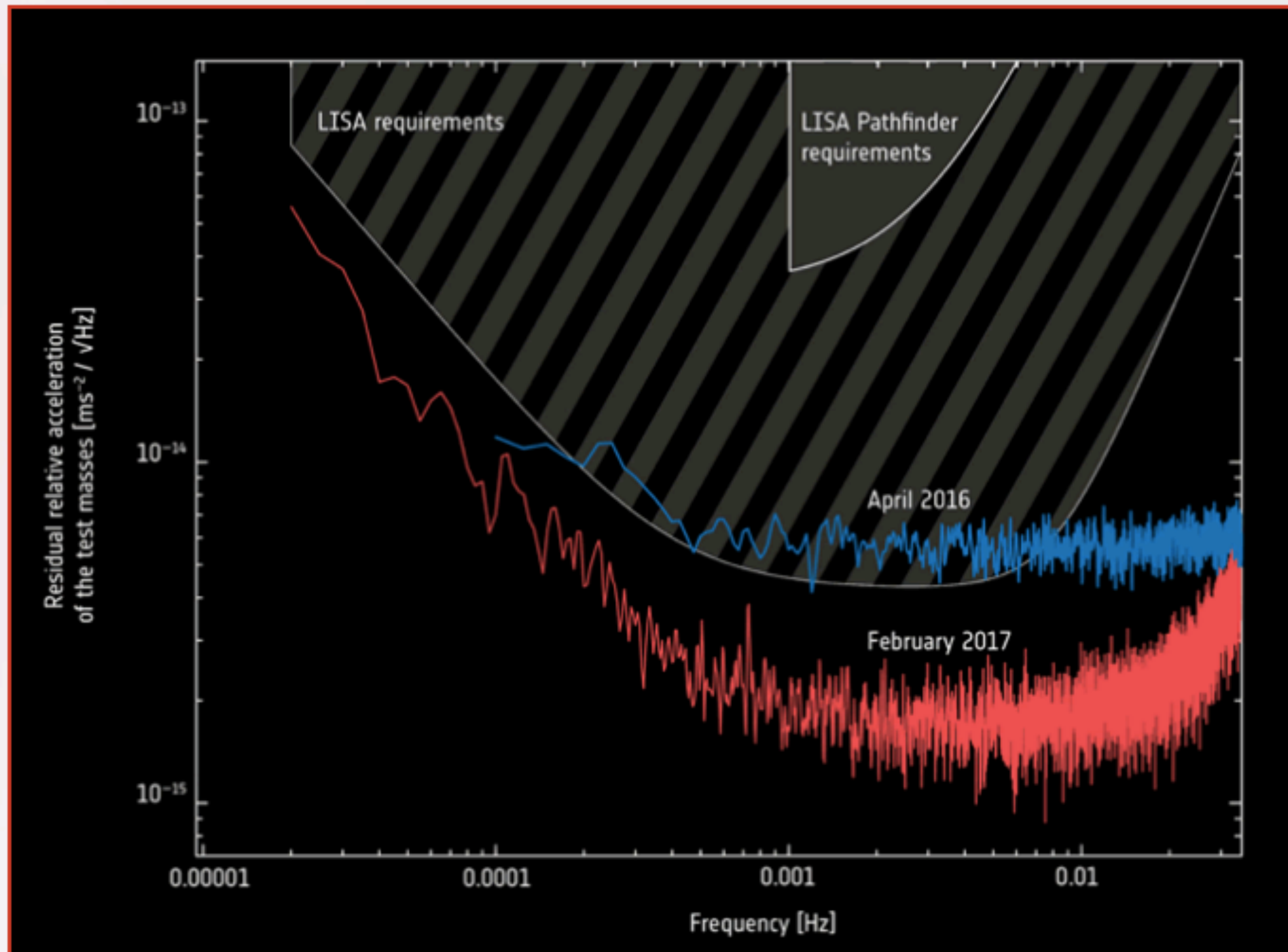
MEASUREMENT PRINCIPLE

- Probe the change in proper time between pairs of **free-falling** test masses caused by GWs
- Proper time is inferred by the time of flight of photons exchanged between the satellites
- We have **multiple links** from which we can form Michelson-like signals

LISA PATHFINDER

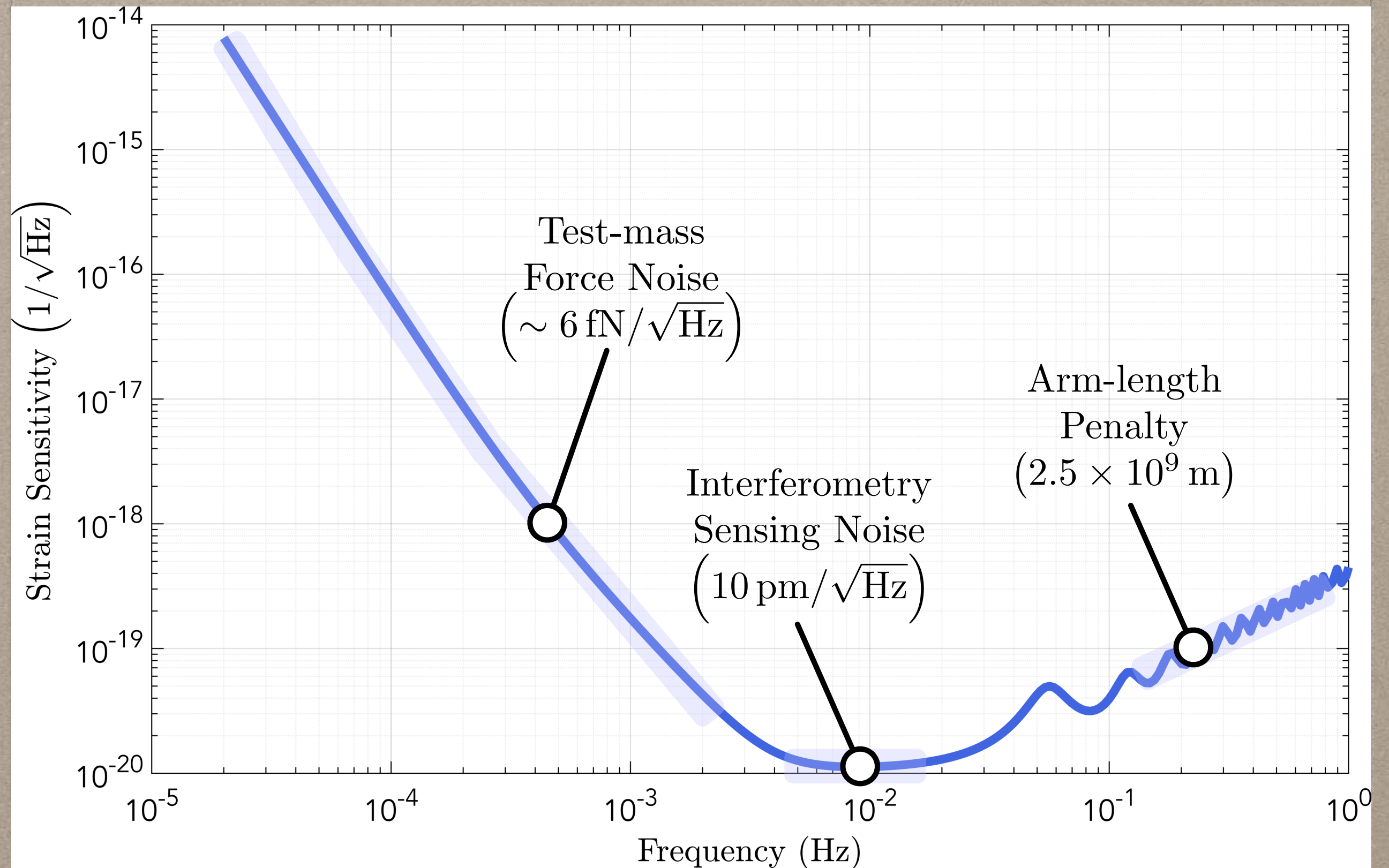


Click here to zoom. LISA will observe a passing gravitational wave directly by measuring the tiny changes in distance between freely falling test masses inside spacecraft with its high precision measurement system. Credit: NASA/ESA



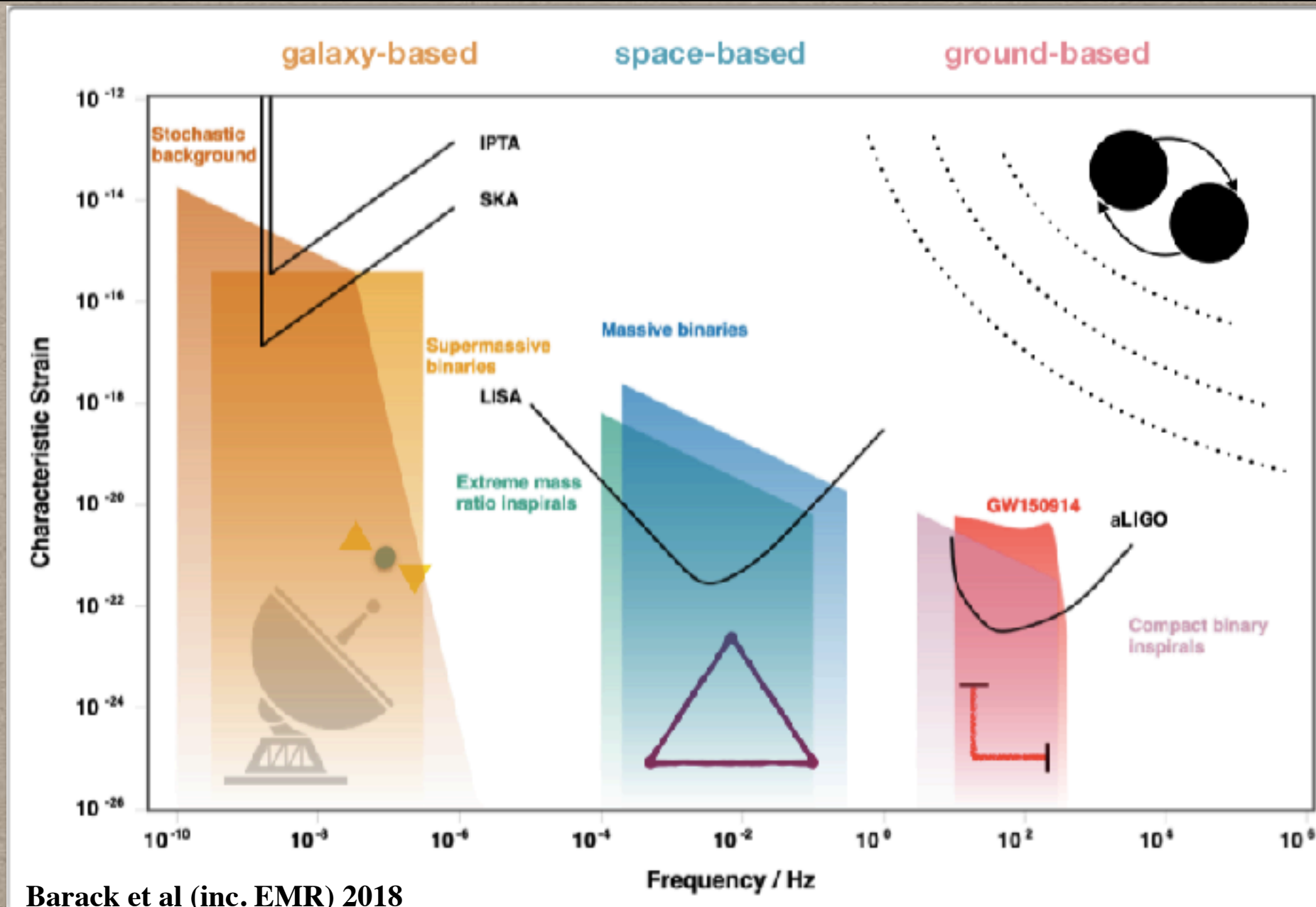
CREDIT: NASA

LISA LIMITING NOISE



LISA IS AN OBSERVATORY

*exploring the 2 mHz ~ 15 minute time domain
for Astrophysics, Cosmology
and Fundamental physics*



SUPERMASSIVE BLACK HOLES (THE LOUDEST)

*Hierarchical growth of structure
in the Universe,
leads us to imply
the existence of SMBH mergers*

(Lacy & Cole 1993)

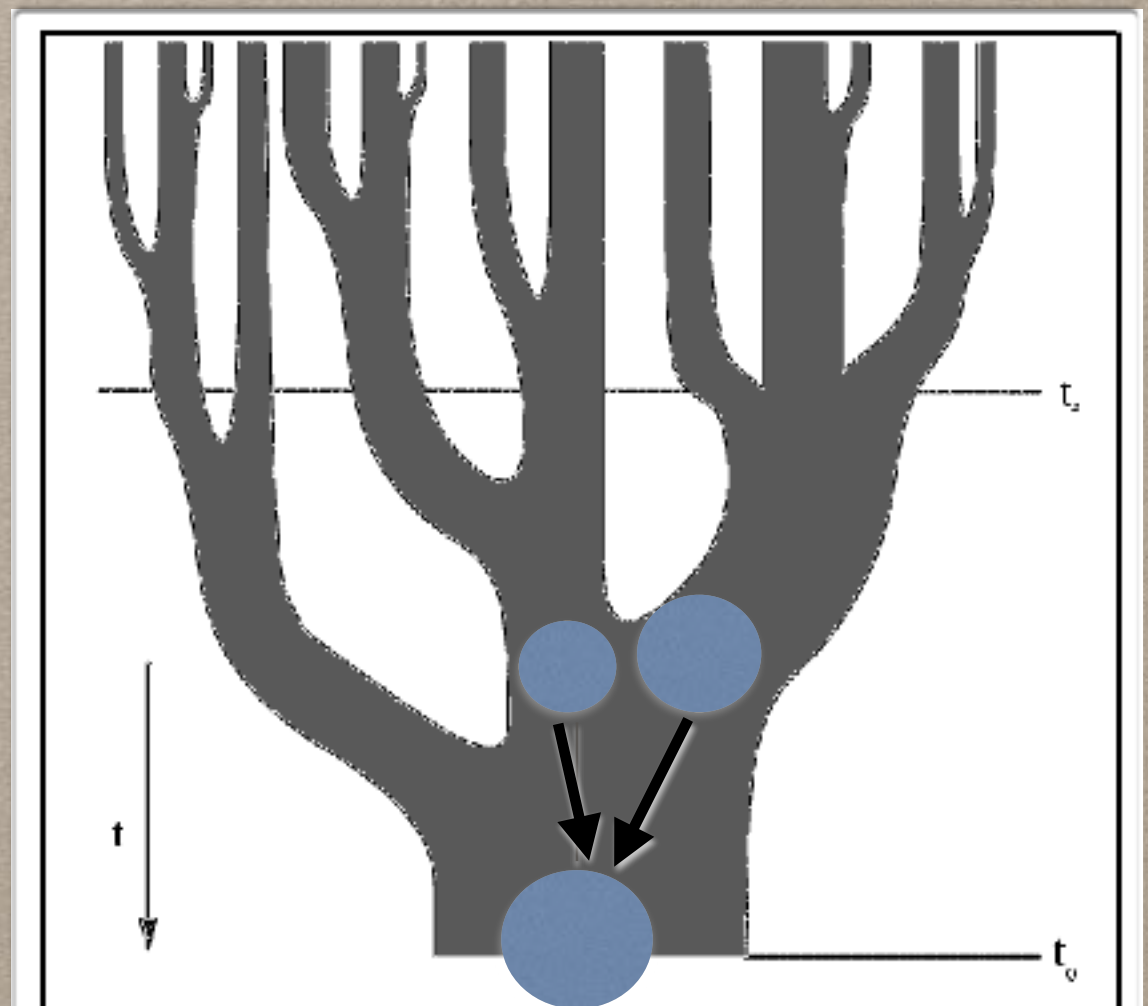
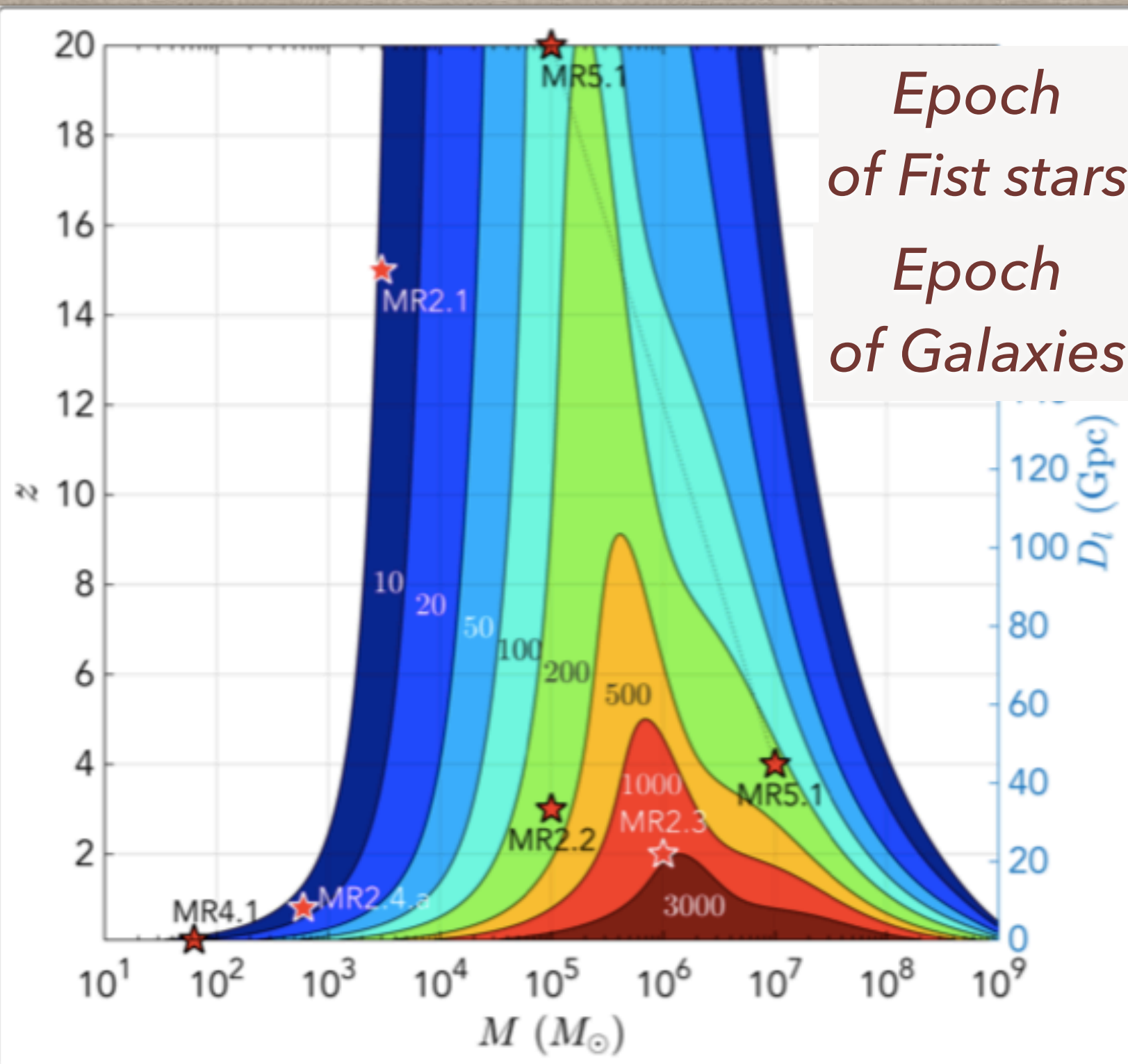


Figure 6. A schematic representation of a "merger tree" depicting the growth of a halo as the result of a series of mergers. Time increases from top to bottom in this figure and the widths of the branches of the tree represent the masses of the individual parent halos. Slicing through the tree horizontally gives the distribution of masses in the parent halos at a given time. The present time t_0 and the formation time t_f are marked by horizontal lines, where the formation time is defined as the time at which a parent halo containing in excess of half of the mass of the final halo was first created.

SUPERMASSIVE BLACK HOLES



*Epoch
of First stars*

*Epoch
of Galaxies*

*LISA is ideal
for studying of
structure formation
beyond the
re-ionisation epoch*

HOW DO SUPERMASSIVE BLACK HOLE FORM ?

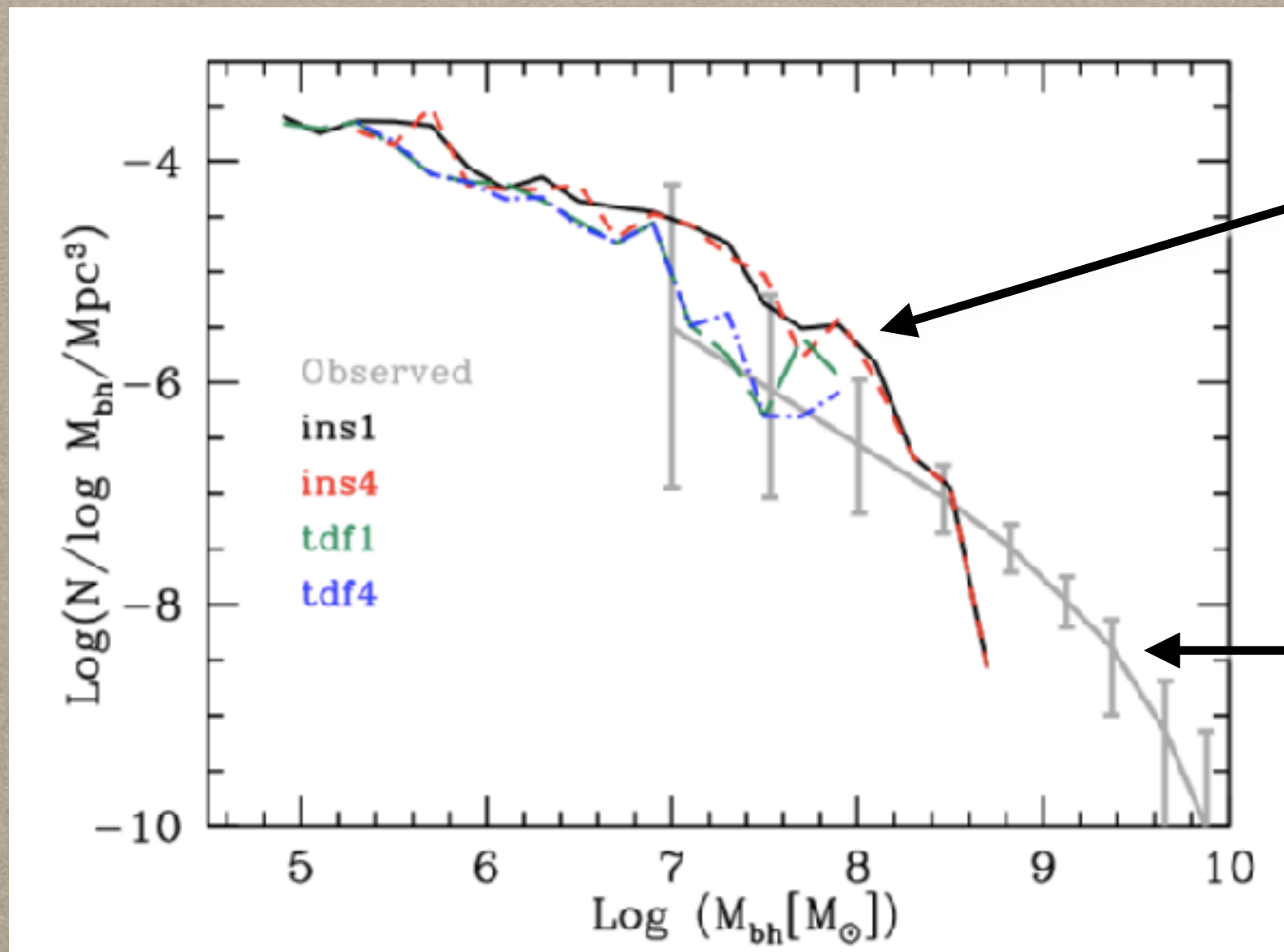
LISA detections to discriminate between formation scenarios:

Dayal, EMR + 2019
DeGraf & Sijacki 2019
Latif et al. 2019
Bonetti et al. 2019
Ricarte & Natarajan 2018
Hartwig, Agarwal & Regan, 2018
Colpi 2018
....

predictions vary by an order of magnitude between papers...

Sesana, Volonteri & Haardt 2007

HOW DO SUPERMASSIVE BLACK HOLE FORM ?

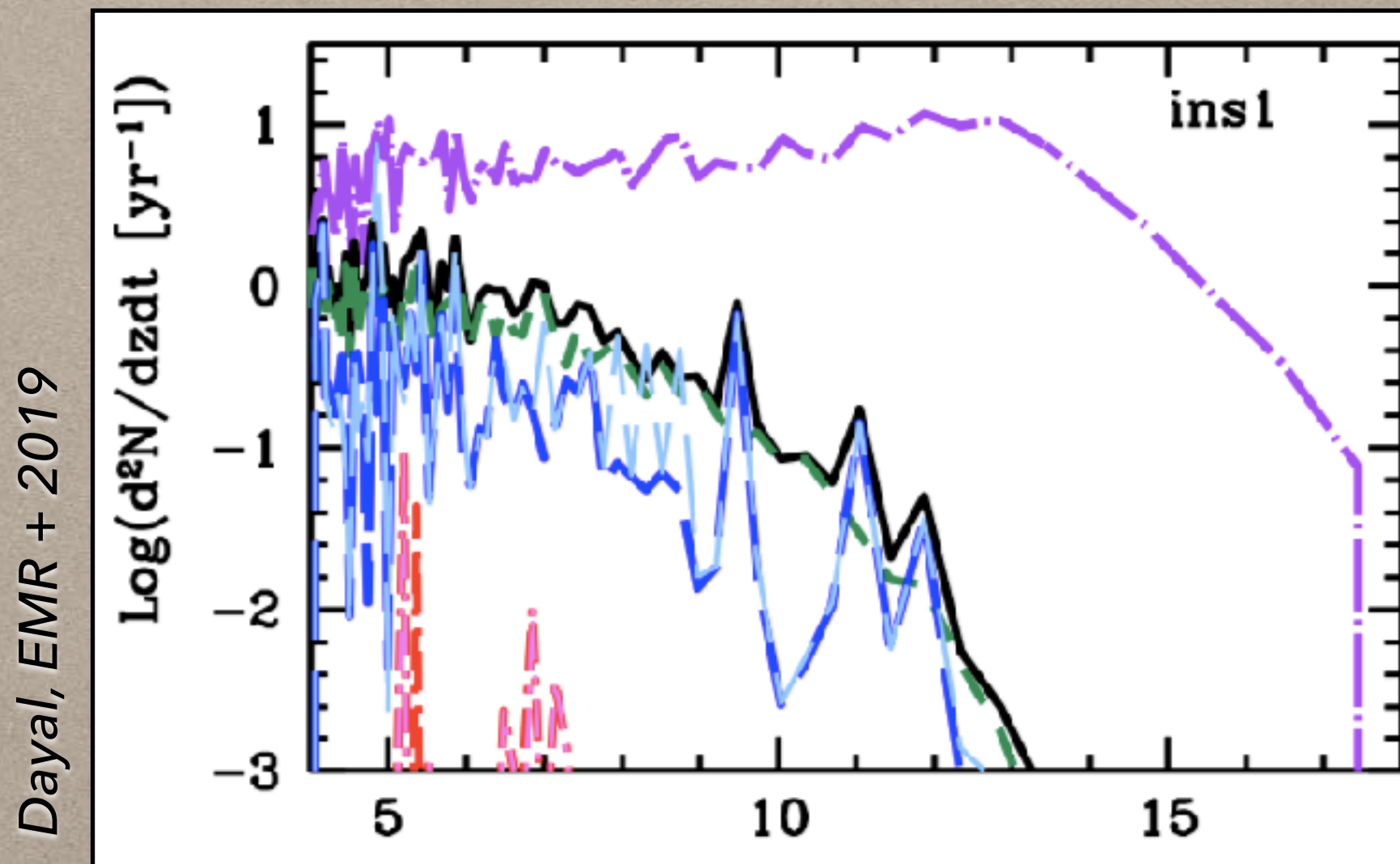


*PoP3 seeds
can describe the
average population*

*need massive
seeds for these*

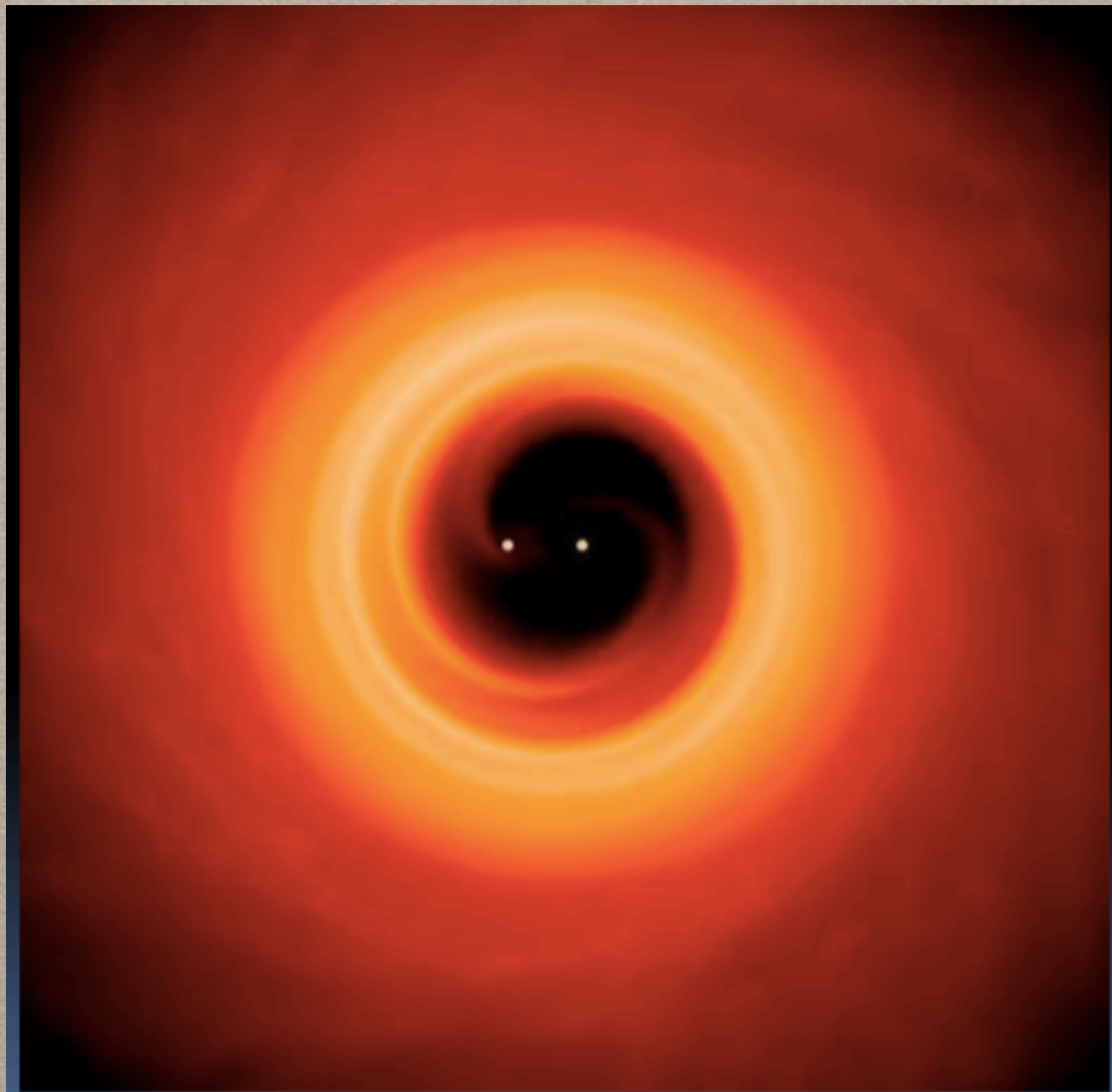
Dayal, EMR + 2019; observation from Willott +2010

A FEW DETECTED PER YEAR...



Most of the information is in the background...

ELECTROMAGNETIC COUNTERPARTS TO MERGER

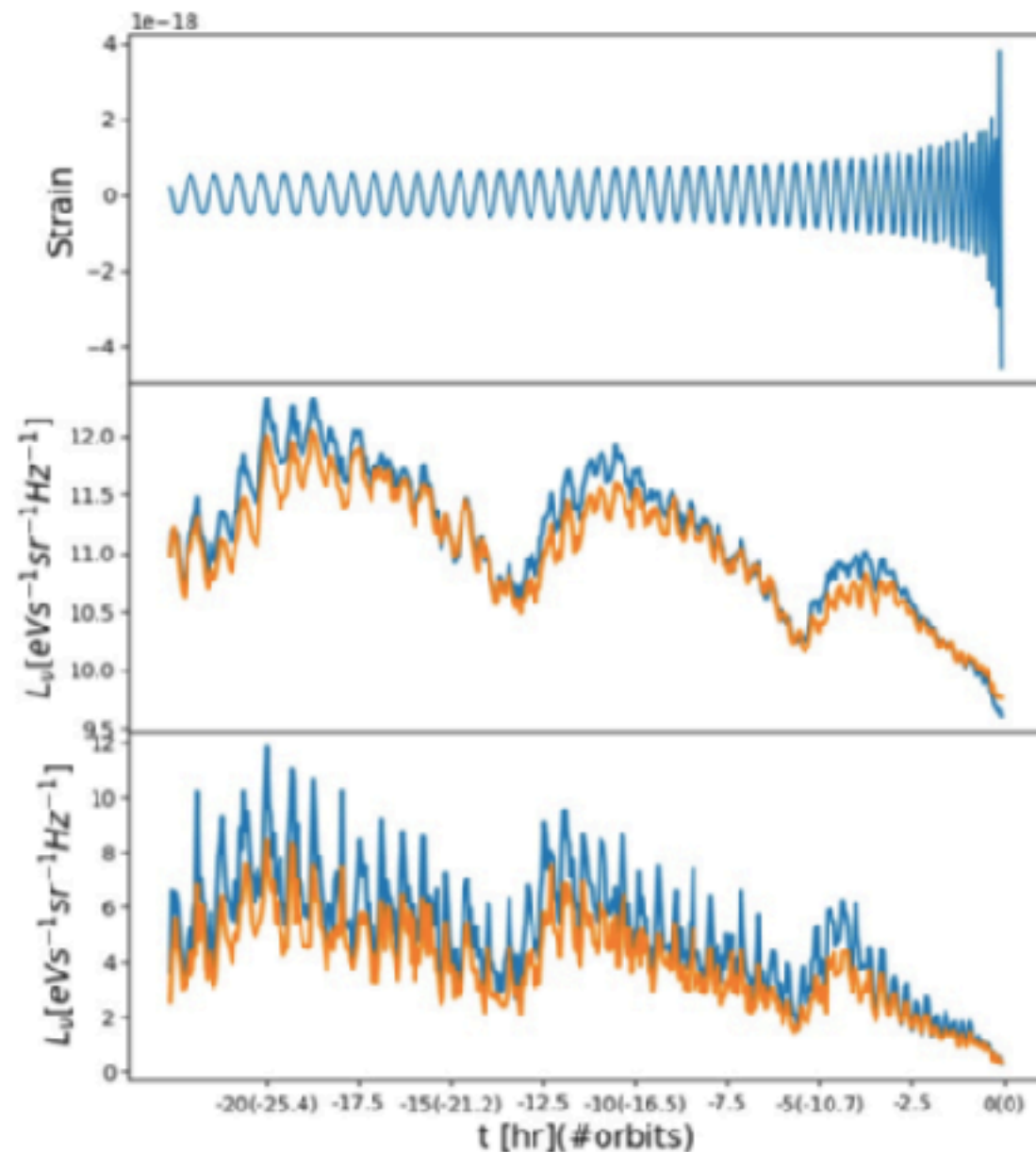
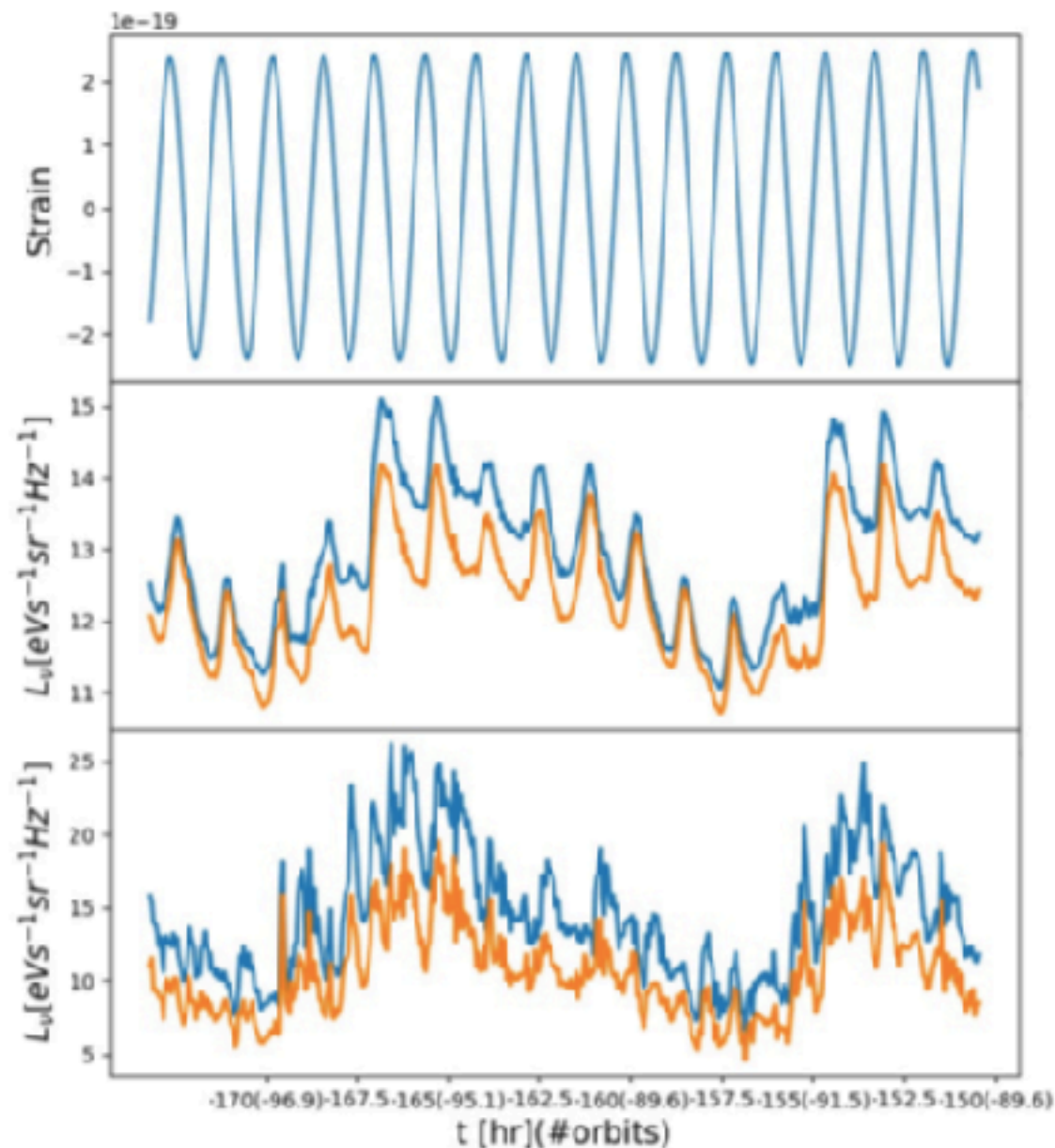


credit J. Cuadra

PRECURSORS TO MERGER

~1 week to merger

~last 2 hr to merger

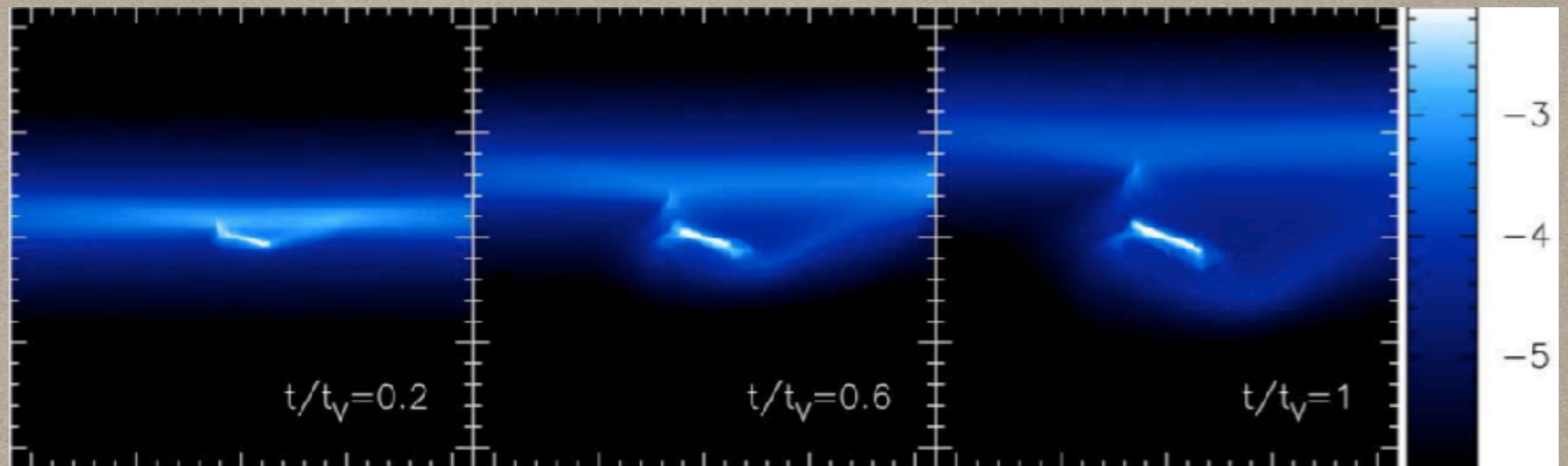


Gravitational
signal

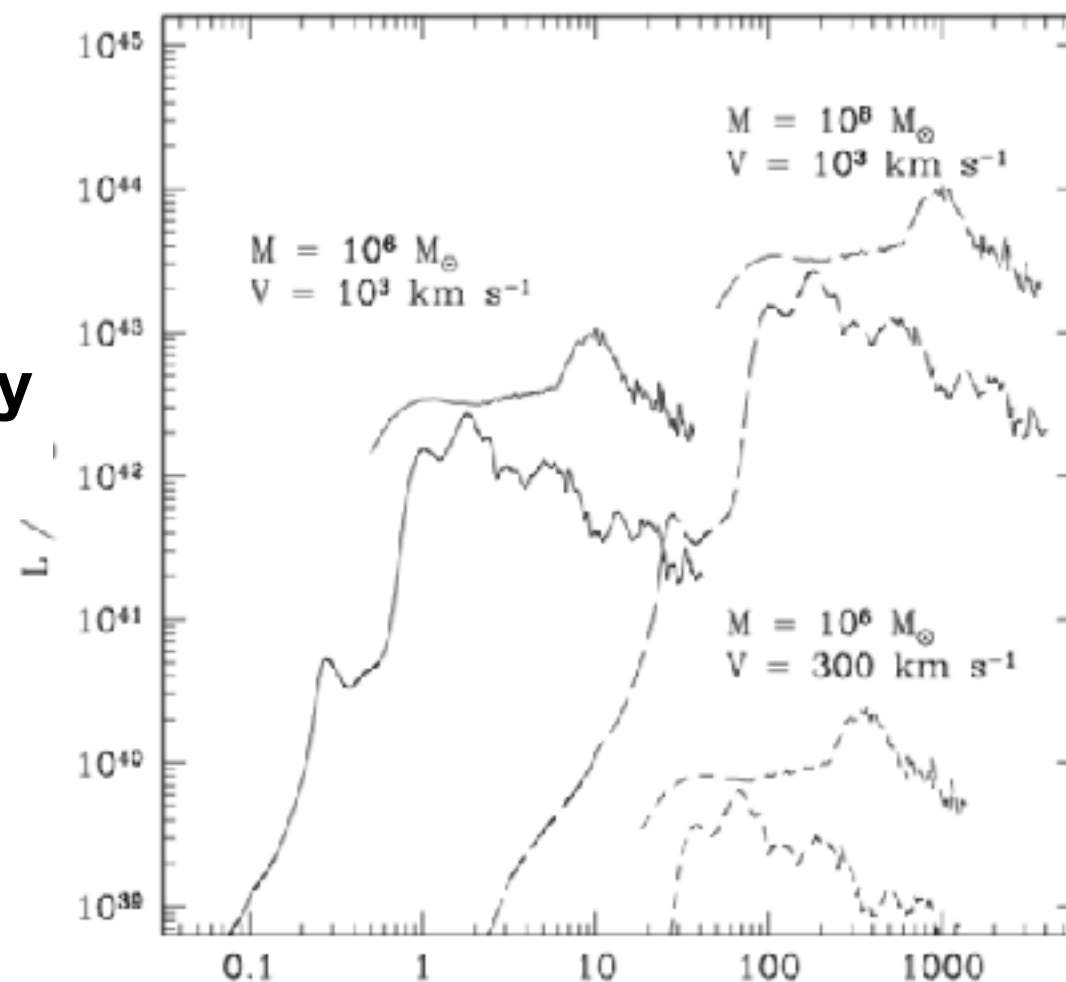
2 KeV

10 KeV

"AFTERGLOWS"



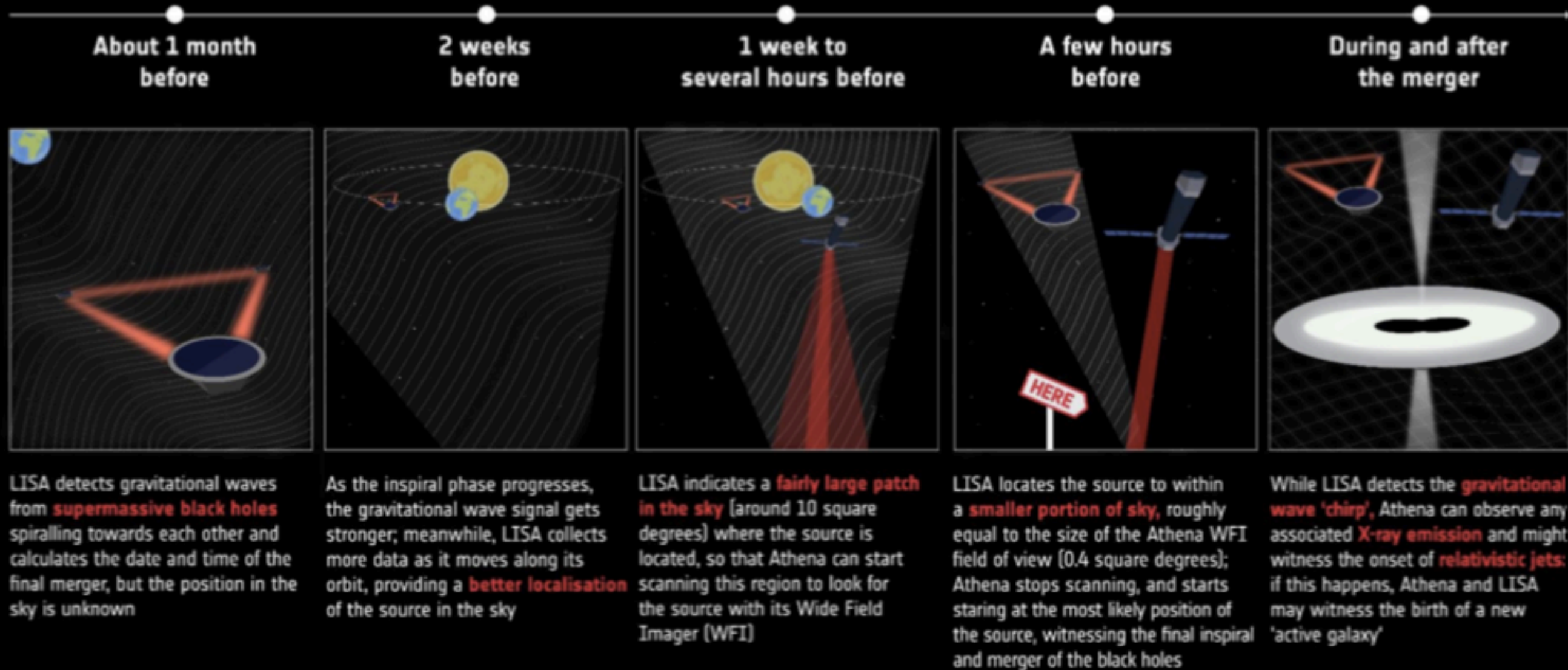
**Luminosity
erg/s**



time (yr)

- Months timescale rise
- Bolometric luminosity at fraction of Eddington

→ HOW CAN LISA AND ATHENA WORK TOGETHER?



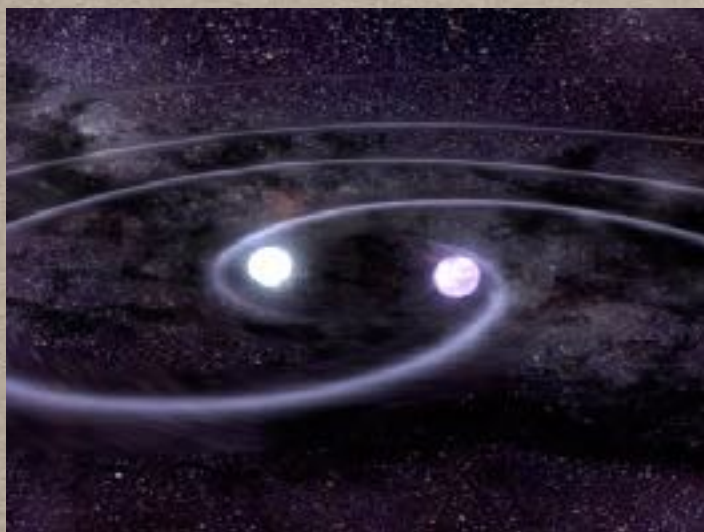
#Space19plus #AnsweringTheBigQuestions

Space19

Courtesy of Paul McNamara

ULTRA COMPACT STELLAR MASS BINARIES (THE MOST NUMEROUS)

**Detached double
white dwarfs**



AM CVn stars



Hot subdwarfs



Predicted:

10^8

10^5

10^8

**Observed
with EM**

~ 10

10^2

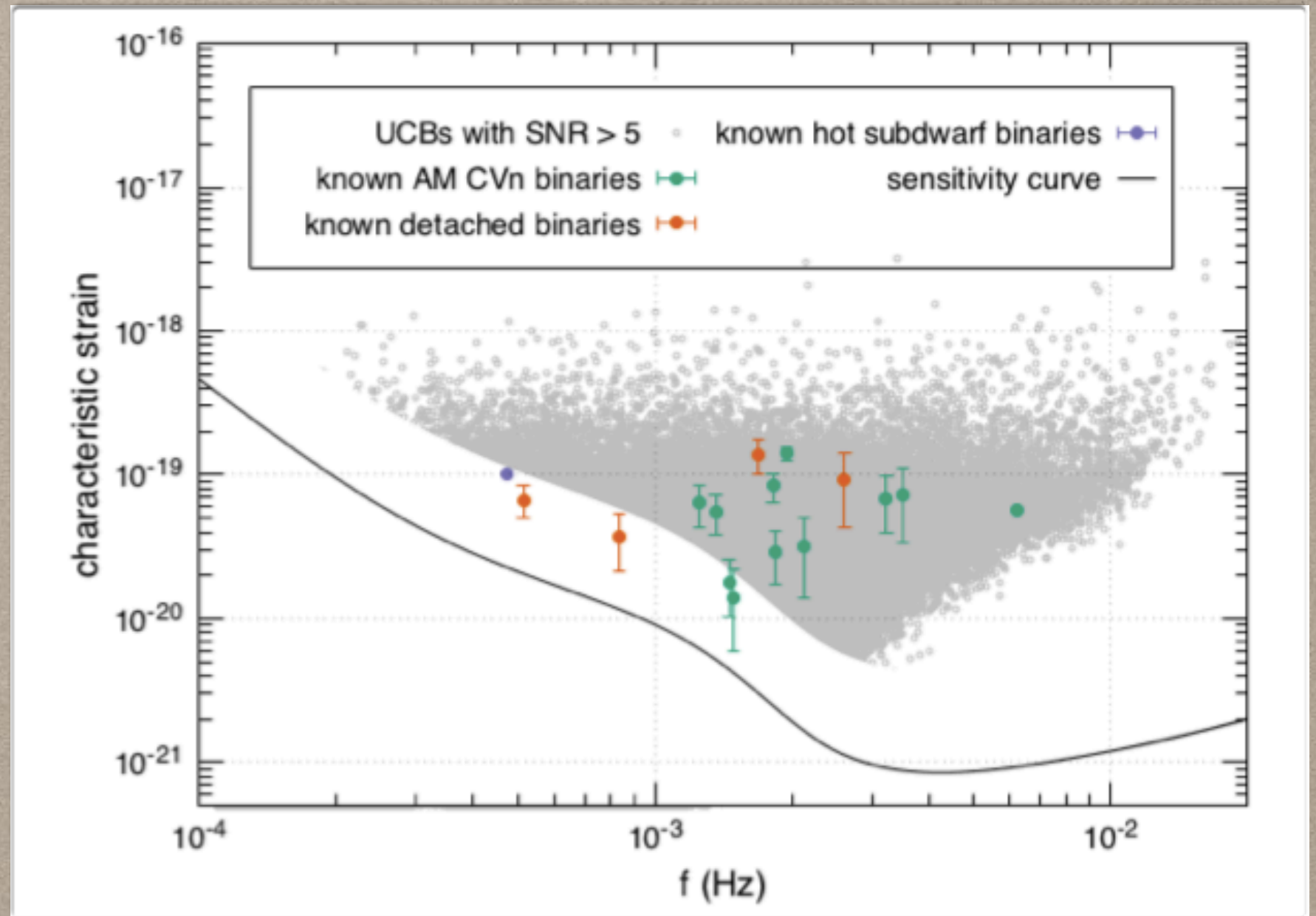
10^3

*see e.g. Astro2020 Science White Paper
Littenberg et al. 2019*

CURRENT SAMPLE

- ~ 10 $SNR > 20$
- 3 $SNR > 100$

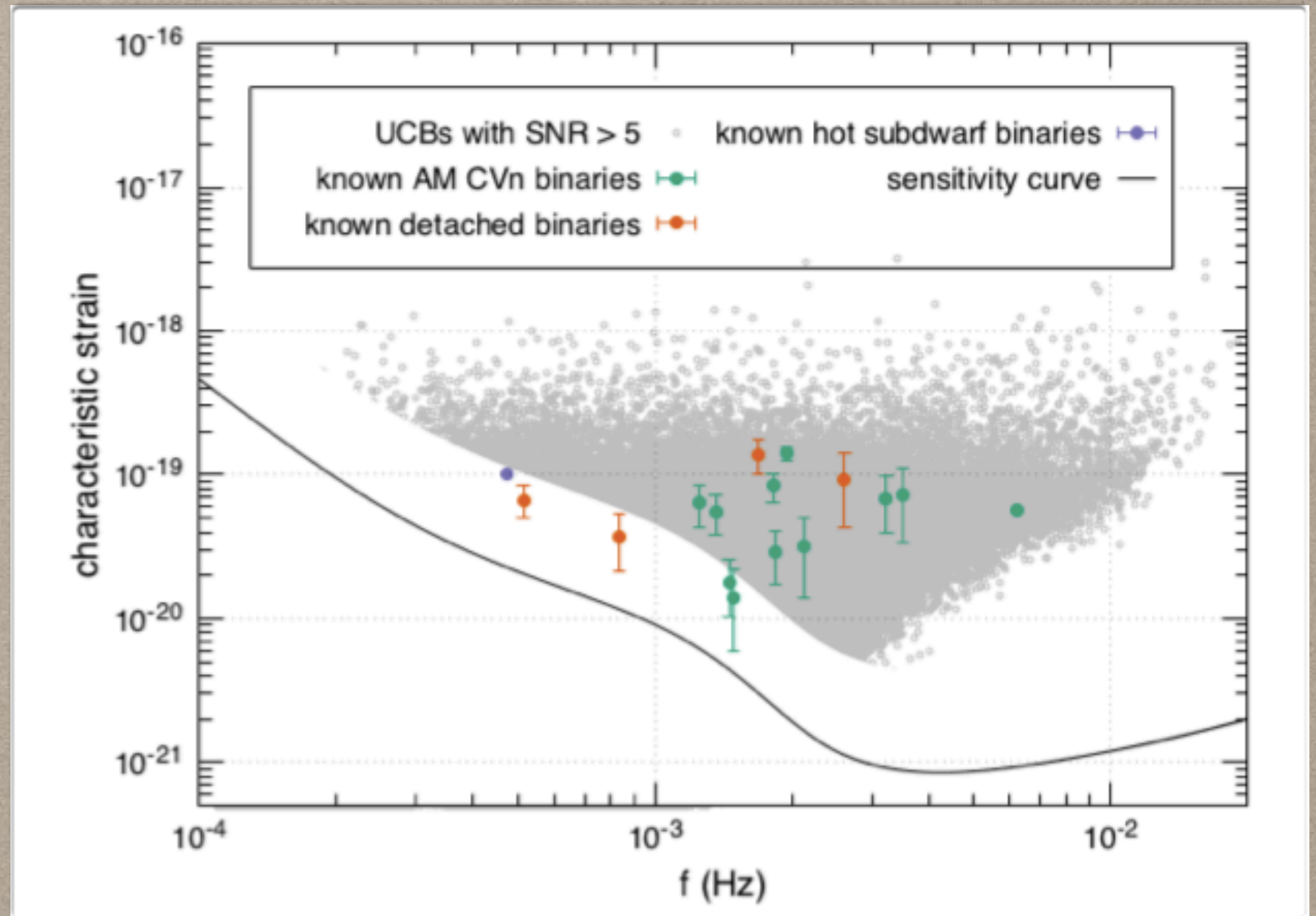
Kupfer, Korol..EMR 2018



Littenberg + 2019

THE FUTURE SAMPLE !

- 25 000 resolved binaries w LISA
- 8 000 precise distance < 0.3
- 5 000 w precise localisation ($\sim \text{arcmin}^2$)
- a few 100 in Gaia and/or LSST



Korol, EMR et al. 2018
Breivik + 2018

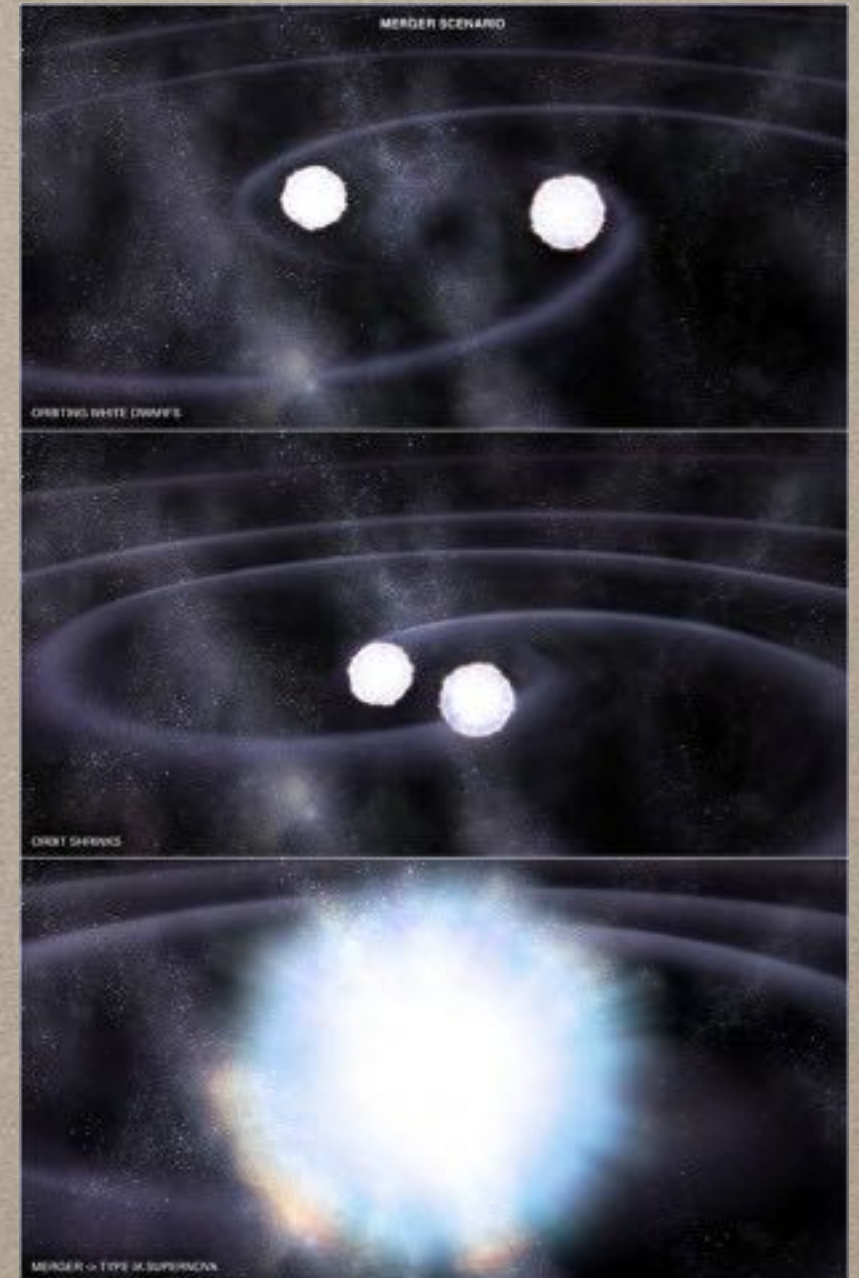
Littenberg + 2019

Gijs + 01, 04

QUEST FOR SUPERNOVA 1A PROGENITORS

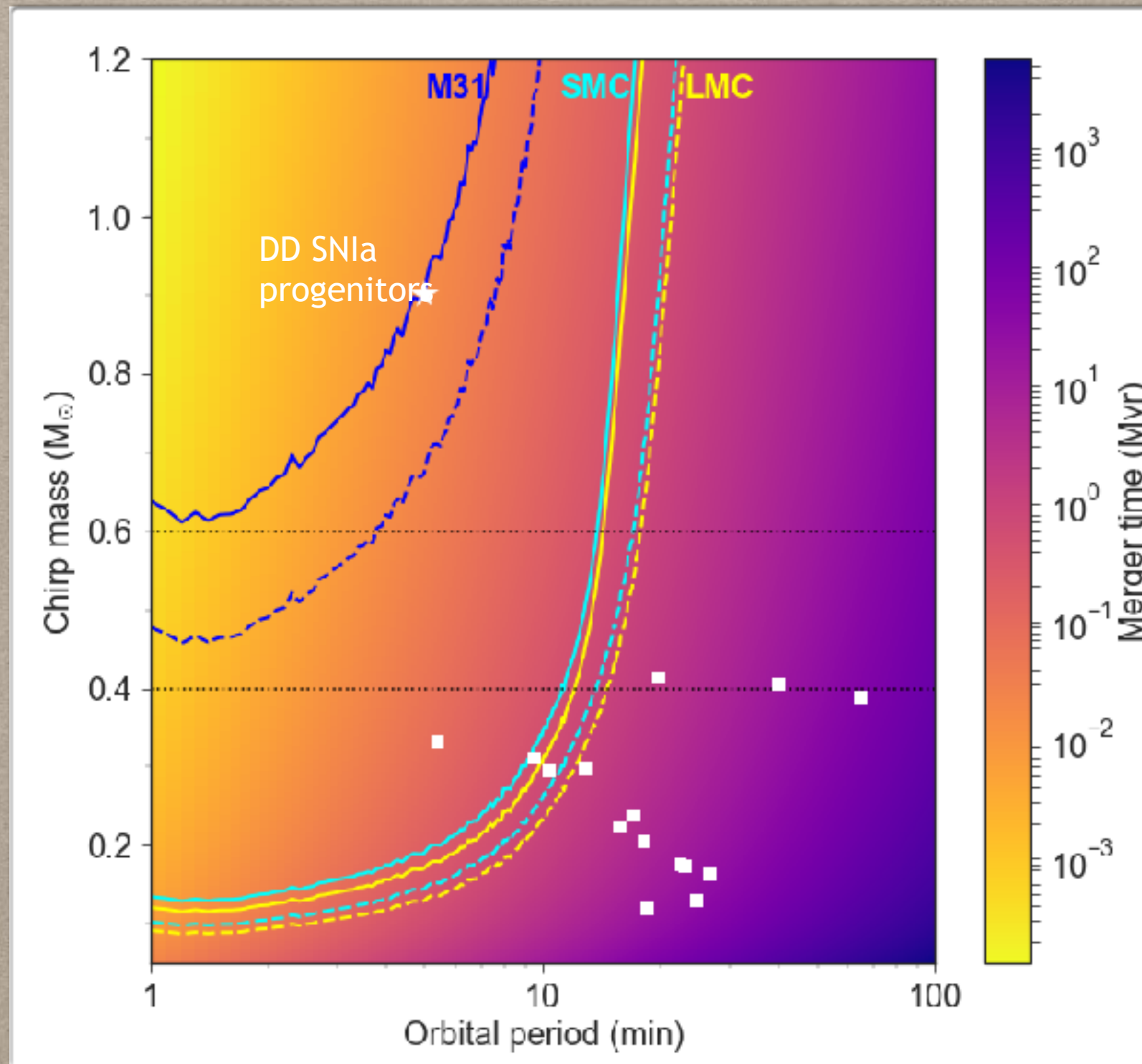
WD-WD binary "merger"

***Massive** system are required but
none has been so far
unambiguously identified in optical
Rebassa-Mansergas + 2018*



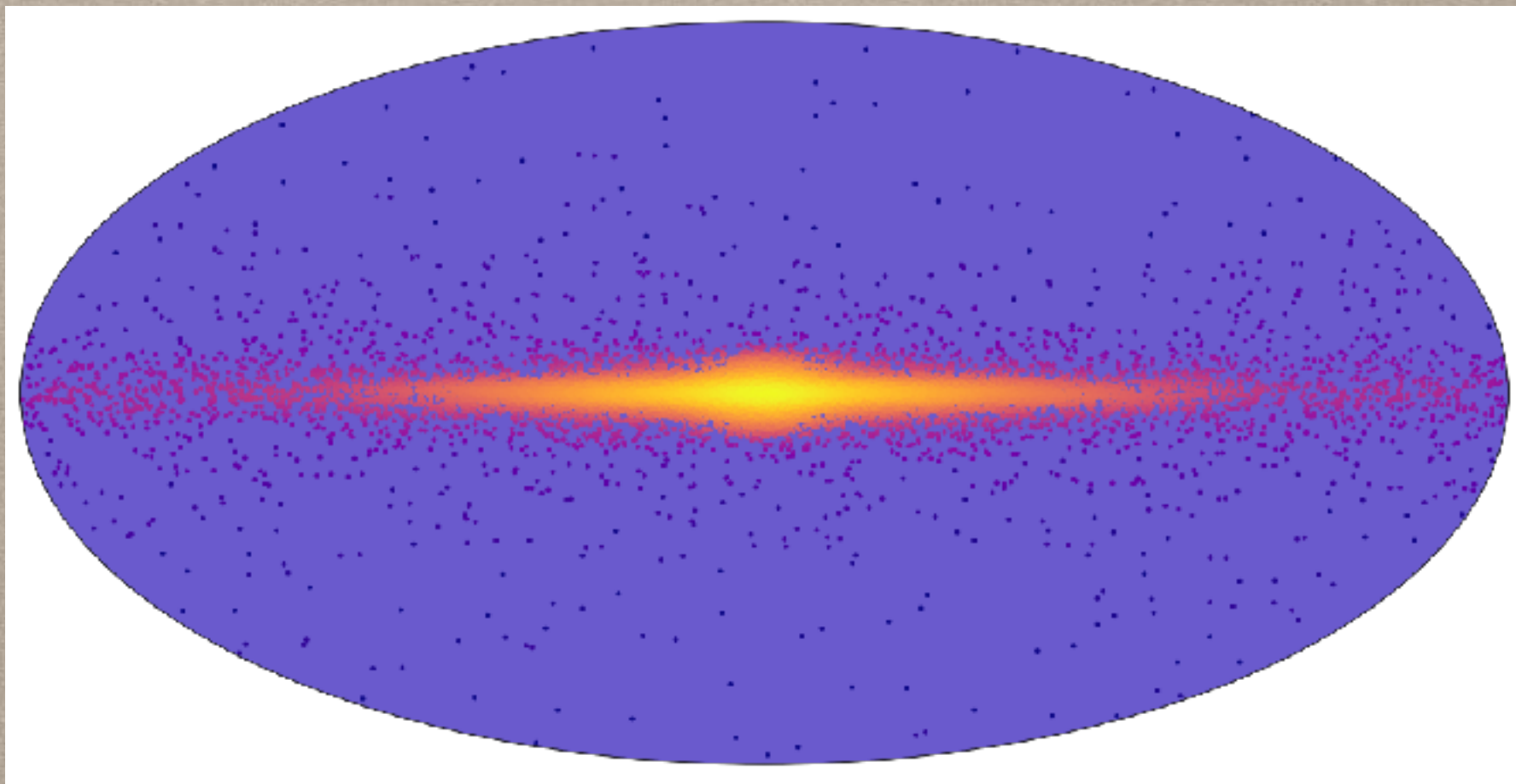
Credit: NASA/CXC/SAO

LISA a unique tool to access SNIa progenitors



NEAR FIELD COSMOLOGY WITH GRAVITATIONAL WAVES

LISA's view of the Milky Way



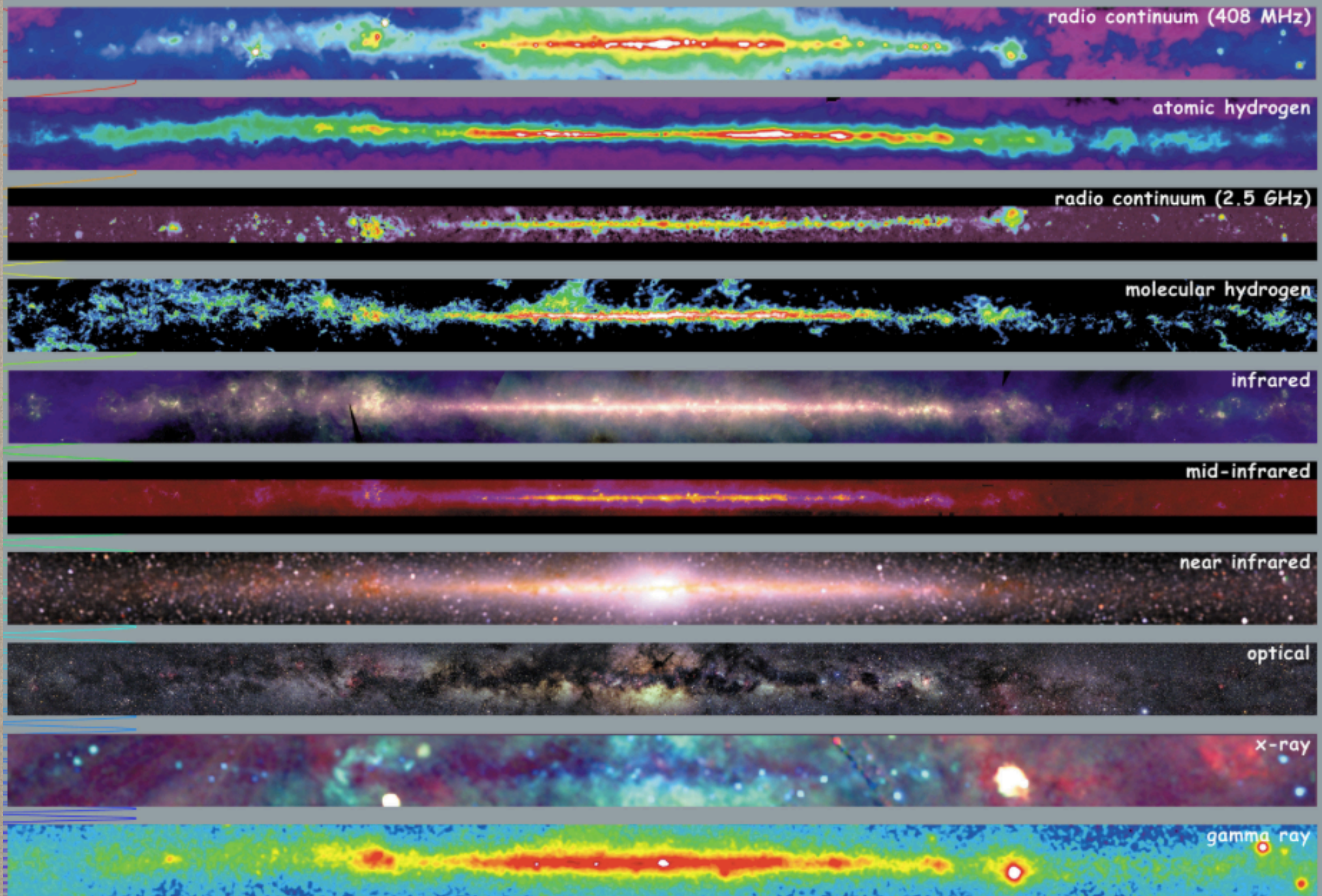
Korol, EMR, Barausse 2018

Lamberts + 2019

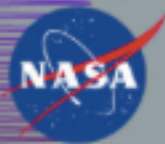
Benacquista & Holley-Bockelmann 2006

Adams, Cornish & Littenberg 2012 , Lambert + 2019

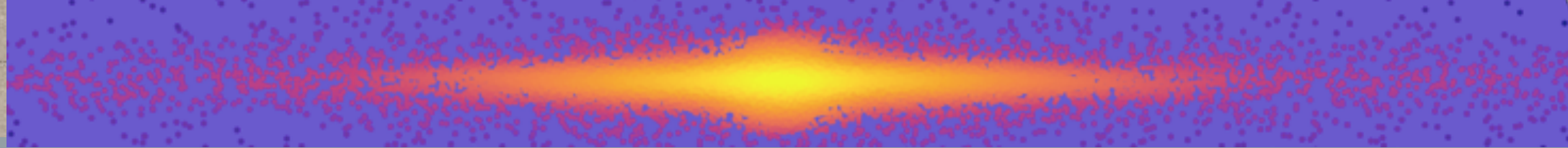
Elena Maria Rossi



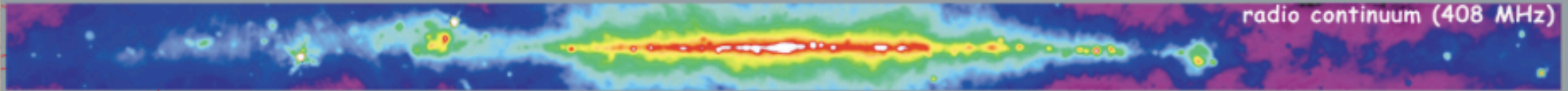
<http://adc.gsfc.nasa.gov/mw>



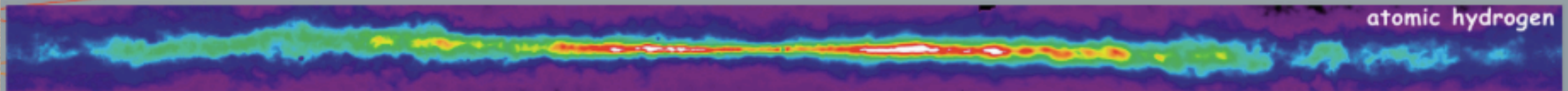
Multiwavelength Milky Way



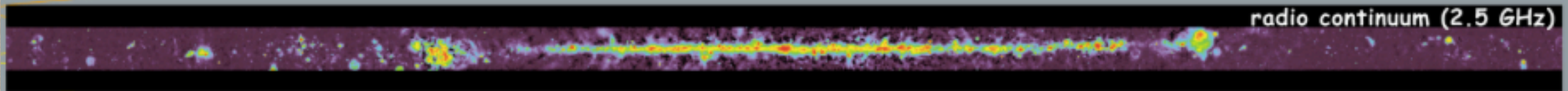
radio continuum (408 MHz)



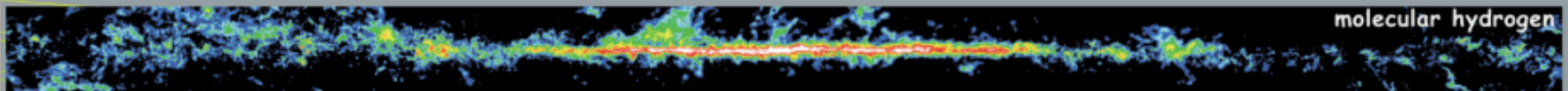
atomic hydrogen



radio continuum (2.5 GHz)



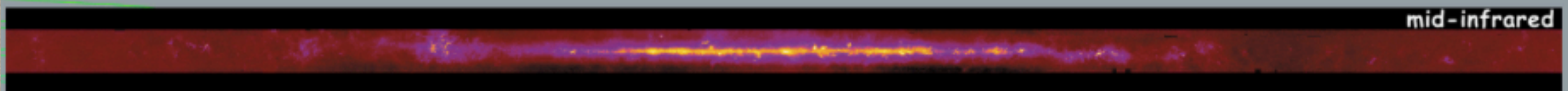
molecular hydrogen



infrared



mid-infrared



near infrared



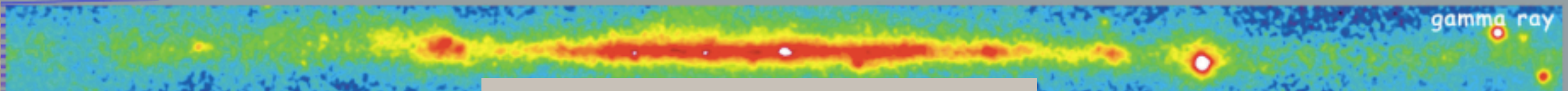
optical



x-ray



gamma ray



<http://adc.gsfc.nasa.gov/mw>



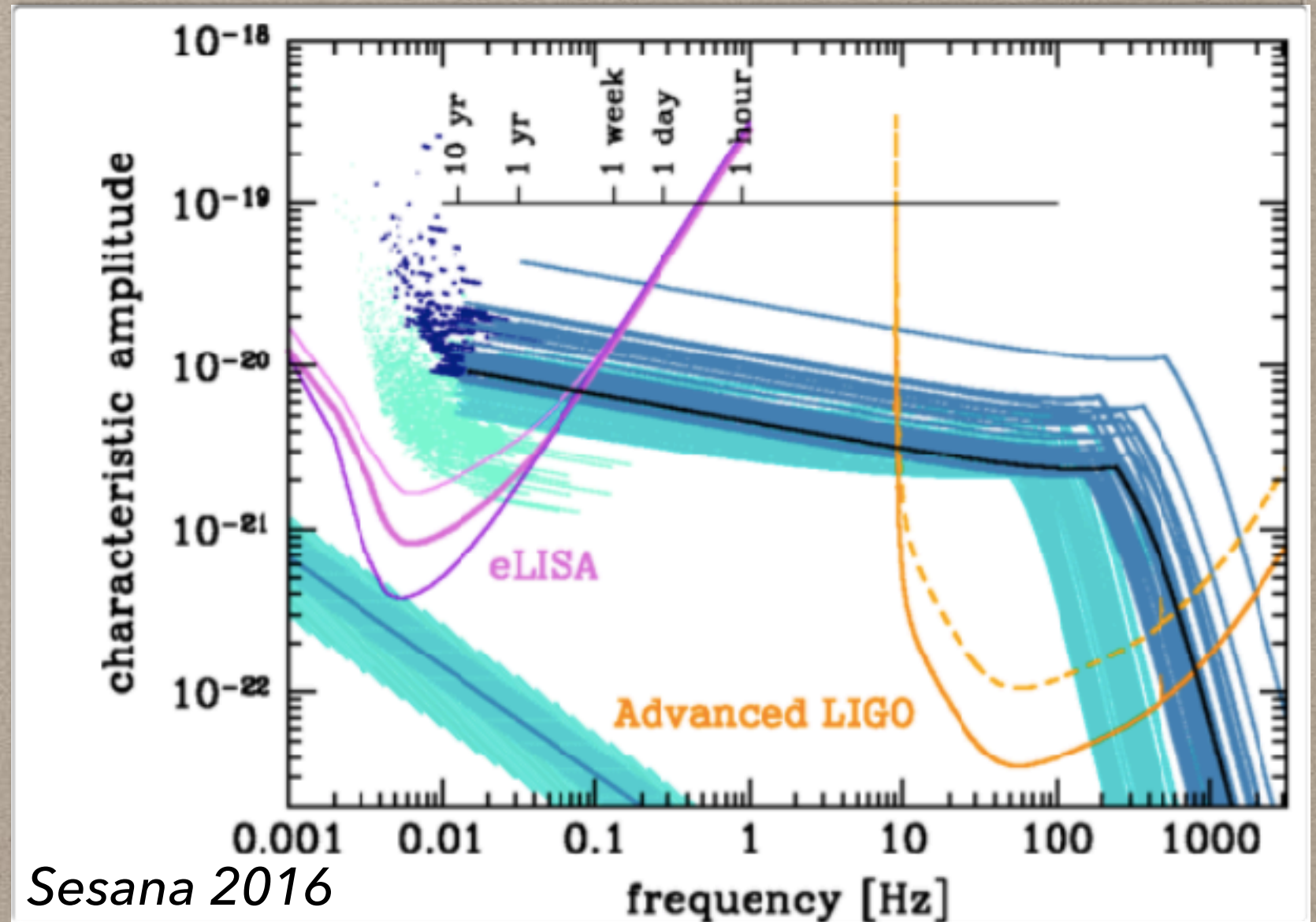
Multi

Messenger

Milky Way

MASSIVE STELLAR MASS BLACK HOLES (A LA LIGO/VIRGO)

- ~100 BHs localised weeks in advance with $< 10\text{s}$ & $< 1\text{ deg}^2$ before merger
- Measuring binary properties (e.g. mass $< 1\%$)
- Searching counterpart when in LISA



Multi-band astrophysics!

FUNDAMENTAL PHYSICS

*Probing dark matter , Mass of Graviton and Test of
General Relativity*

*see e.g. Astro2020 Science White Paper (Berti +2019)
Gair et al. Living Review , and Barack + 2018*

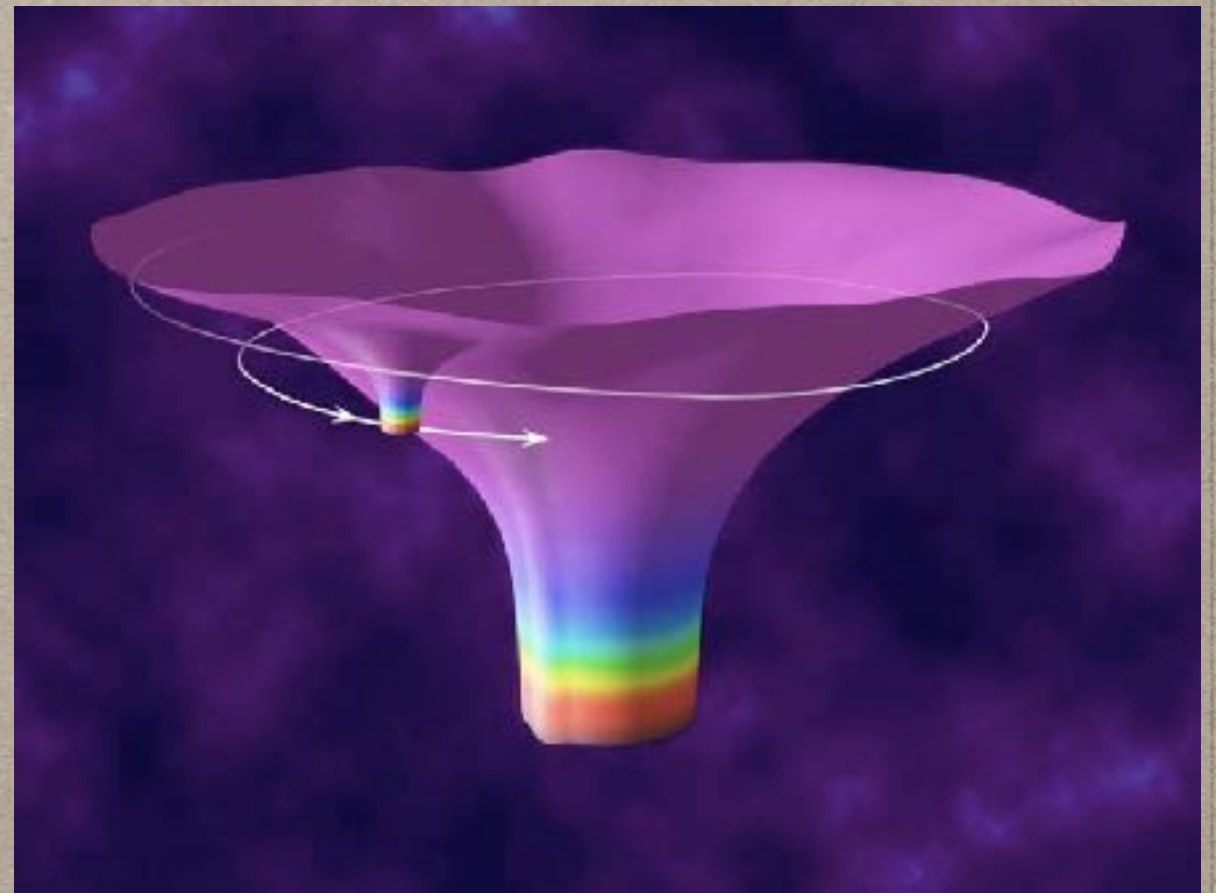
TESTS OF GENERAL RELATIVITY

SMBH-SMBH binary

Extreme Mass Ratio Inspiral



Image credit: The SXs Project



TESTS OF GENERAL RELATIVITY

SMBH-SMBH binary



Image credit: The SXS Project

testing deviation from GR in

- *GW propagation*
- *BH dynamics*

TESTS OF GENERAL RELATIVITY

SMBH-SMBH binary

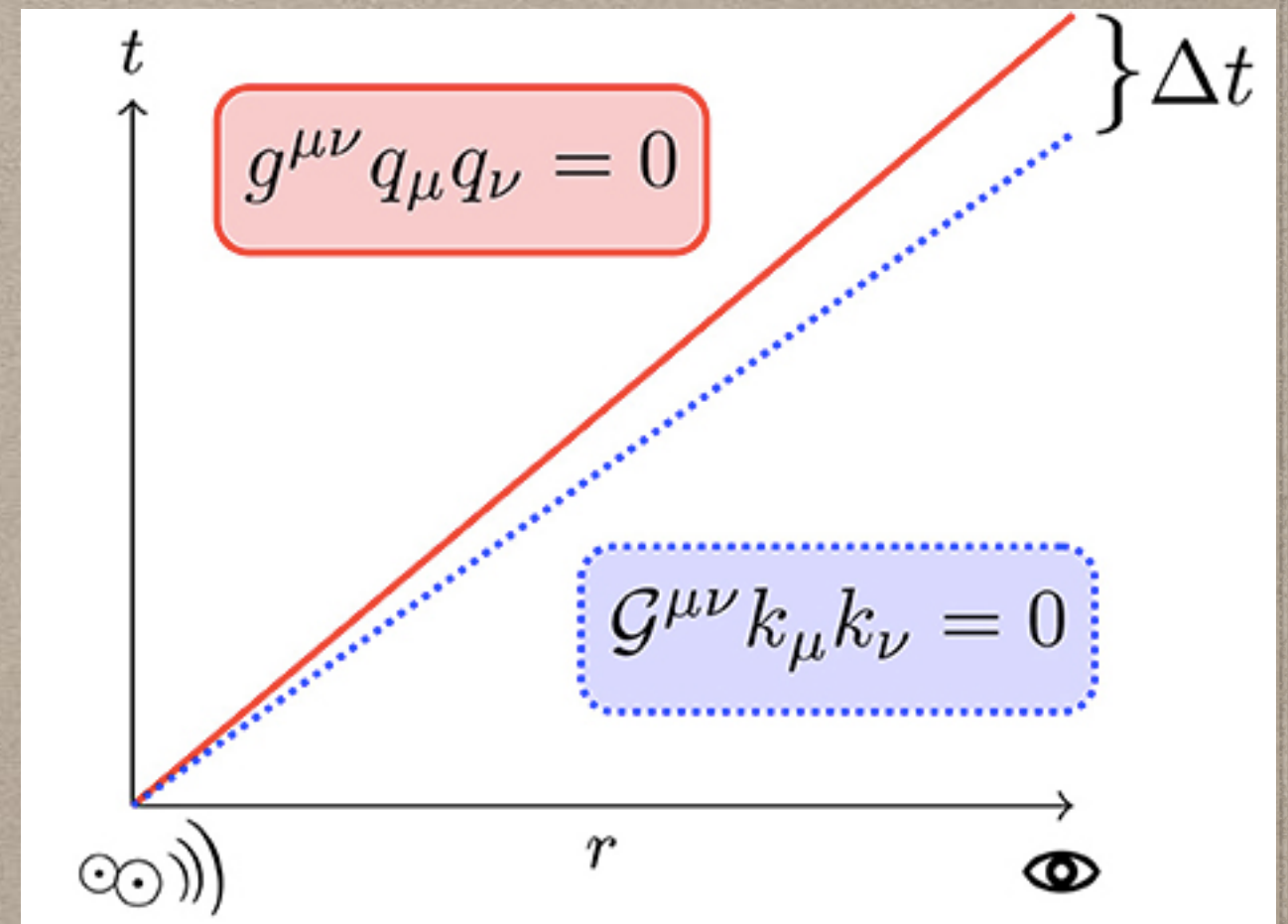
Cartoon by Pilar Ruiz-Lapuente



Image credit: The SX5 Project

testing deviation from GR in

- **GW propagation**
- BH dynamics



*In modified gravity theory
there is a time delay between
photons and GW*

TESTS OF GENERAL RELATIVITY

SMBH-SMBH binary

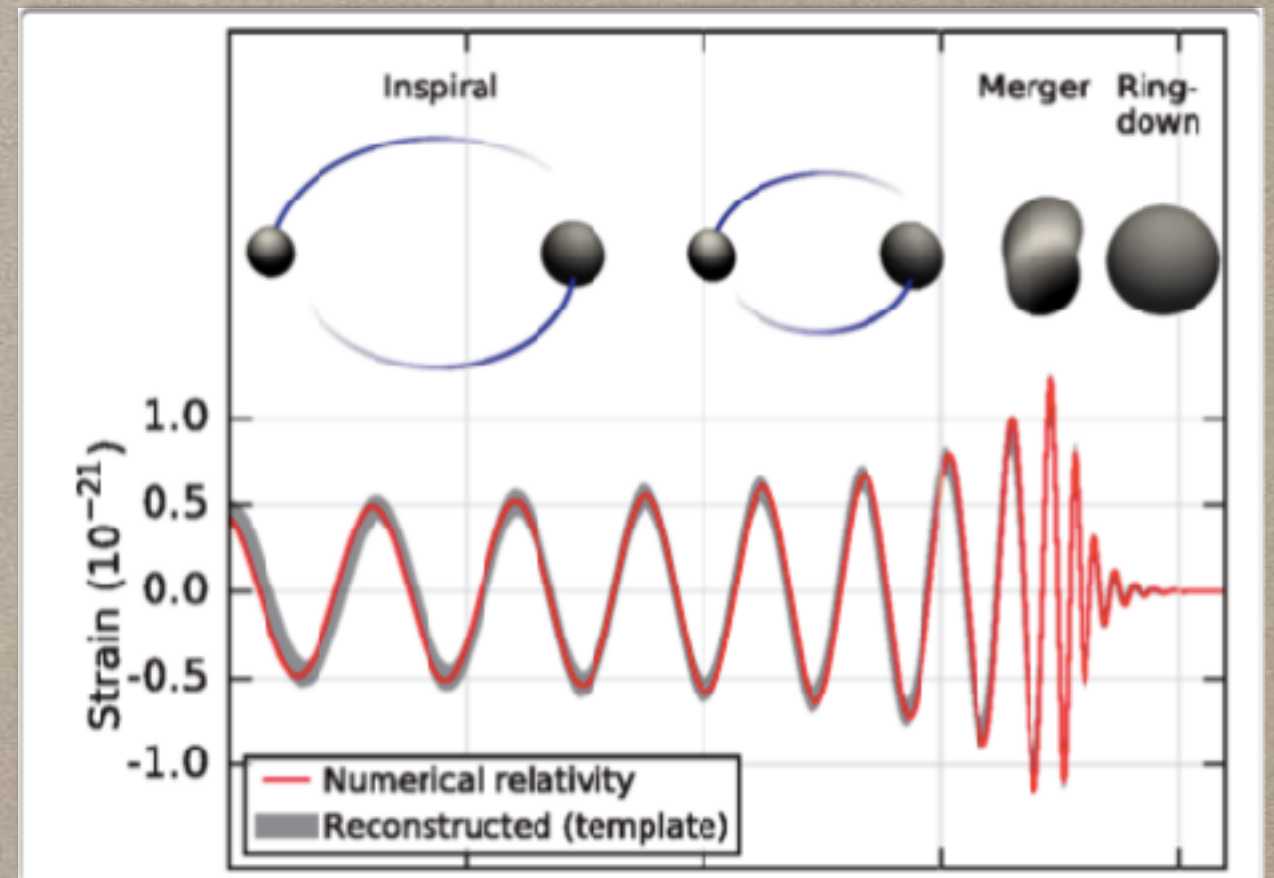
credit: LIGO/VIRGO collaboration



Image credit: The SXS Project

testing deviation from GR in

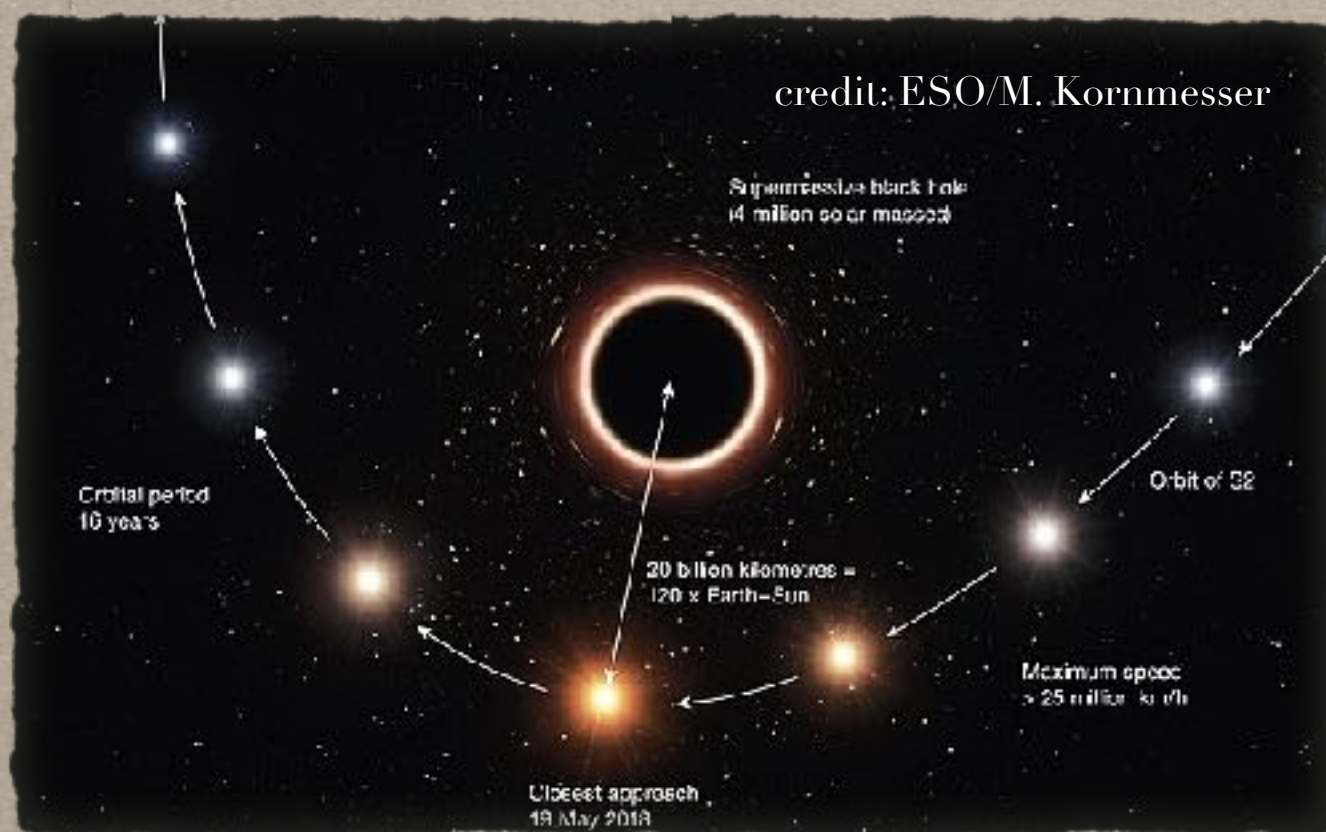
- GW propagation
- **BH dynamics**



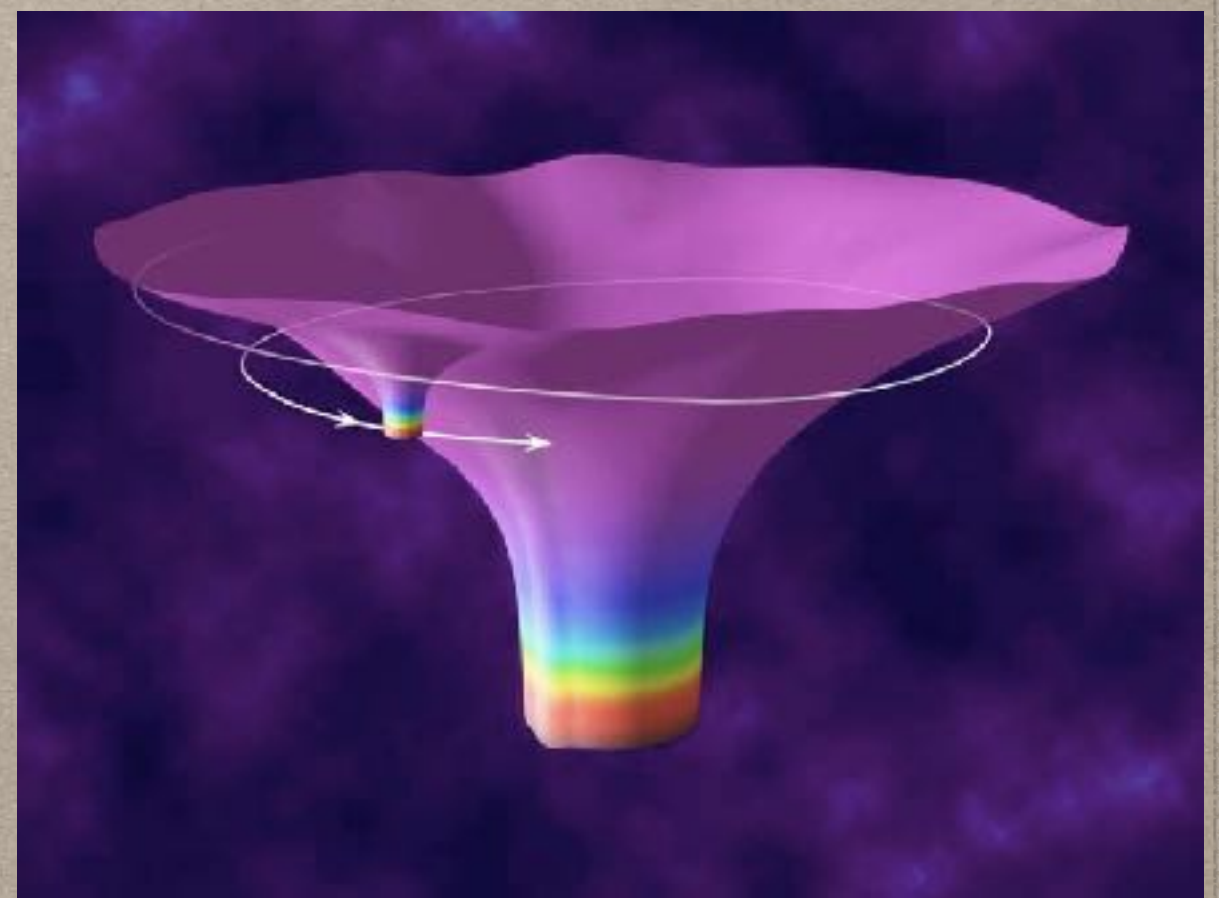
Testing no-hair theorem with
BH spectroscopy in a dynamic,
non-stationary spacetime (Ring-
down phase)

TESTS OF BH SPACETIME

Stars orbiting SgrA* ~120 AU



Extreme Mass Ratio Inspiral



*With 10^4 - 10^5 GW cycles in the LISA band, measurements of deviation from Kerr metric at **0.01-1% level** with a few to a thousands EMRIs, with SNR up to a few 100*

COSMOLOGY

see e.g. Astro2020 Science White Paper Caldwell + 2019

Extremely rich science..

COSMOLOGY

see e.g. *Astro2020 Science White Paper Caldwell + 2019*

- *Measurements of the cosmological parameters using LISA standard sirens (Tamanini + 2017)*
- *Testing modified gravity at cosmological distances with LISA standard sirens (Belgacem +2019)*
- *Testing cosmic strings*
- *Testing particle physics models' predictions for first-order phase transitions*
- *Testing Inflationary models*
- *Cosmological stochastic background (e.g. Caprini +2019)*

MULTIMESSANGER SCIENCE OPPORTUNITIES WITH LISA

Primary Thematic Science Area: Multi-Messenger Astronomy and Astrophysics
Secondary Areas: Cosmology and Fundamental Physics, Galaxy Evolution,
Formation and Evolution of Compact Objects

Multimessenger science opportunities with mHz gravitational waves

John Baker,^{1,2} Zoltán Haiman,³ Elena Maria Rossi,⁴
Edo Berger,⁵ Niel Brandt,⁶ Elmé Breedt,⁷ Katelyn Breivik,⁸ Maria Charisi,⁹ Andrea
Derdzinski,³ Daniel J. D’Orazio,⁵ Saavik Ford,^{10, 11} Jenny E. Greene,¹² J. Colin Hill,^{13, 14}
Kelly Holley-Bockelmann,¹⁵ Joey Shapiro Key,¹⁶ Bence Kocsis,¹⁷ Thomas Kupfer,¹⁸ Shane
Larson,¹⁹ Piero Madau,²⁰ Thomas Marsh,²¹ Barry McKernan,^{10, 11} Sean T. McWilliams,²²
Priyamvada Natarajan,²³ Samaya Nissanke,²⁴ Scott Noble,^{25, 1} E. Sterl Phinney,⁹ Gavin
Ramsay,²⁶ Jeremy Schnittman,¹ Alberto Sesana,^{27, 28} David Shoemaker,²⁹ Nicholas Stone,³
Silvia Toonen,^{30, 27} Benny Trakhtenbrot,³¹ Alexey Vikhlinin,⁵ and Marta Volonteri³²

LISA SCIENCE GROUP

LSG

chairs

WPT: waveforms

WPT:
data-analysis tools

WPT: low-latency
pipelines

WPT: source
identification codes

WPT: catalogs

WPT: multi-
messenger astro

WPT: science
interpretation

simulation WG

astrophysics WG

cosmology WG

waveforms WG

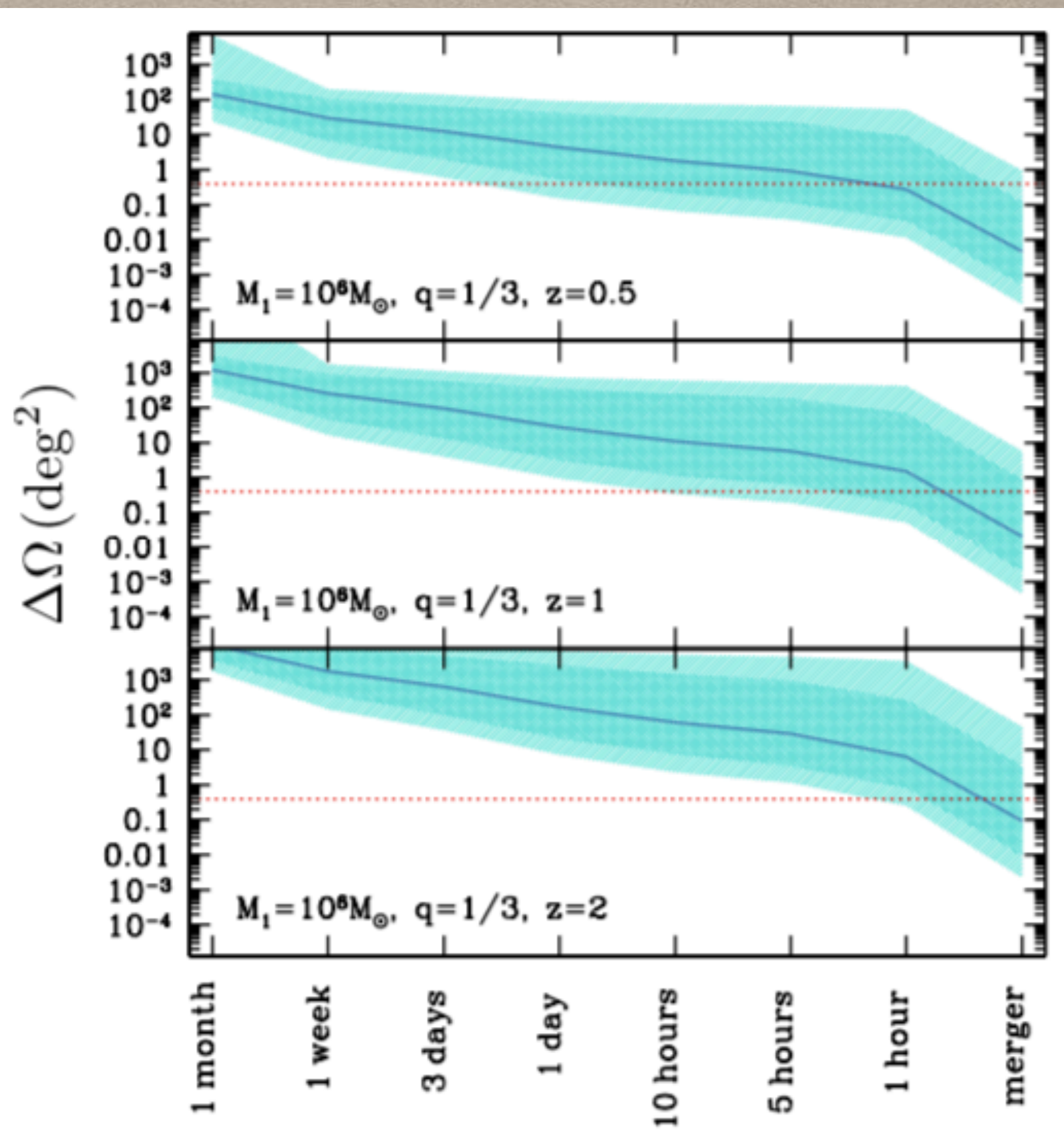
fund. physics WG

LISA data challenges

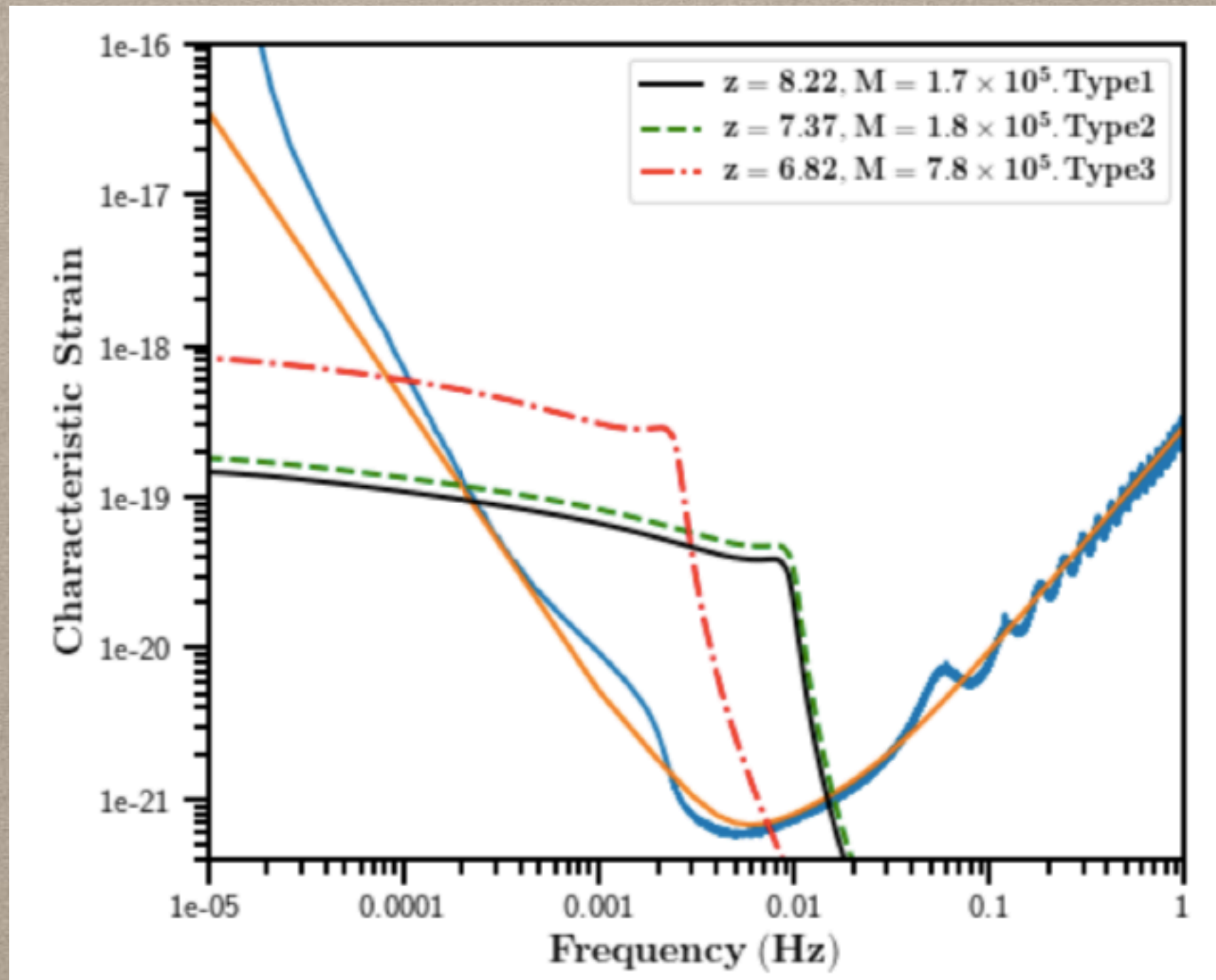
CHAIRS: JON GAIR, MICHELE VALLISNERI & ELENA MARIA ROSSI

BACK UP SLIDES

PRE-MERGER LOCALISATION



DIFFERENT SEEDING MODELS



Dayal, EMR + 2018